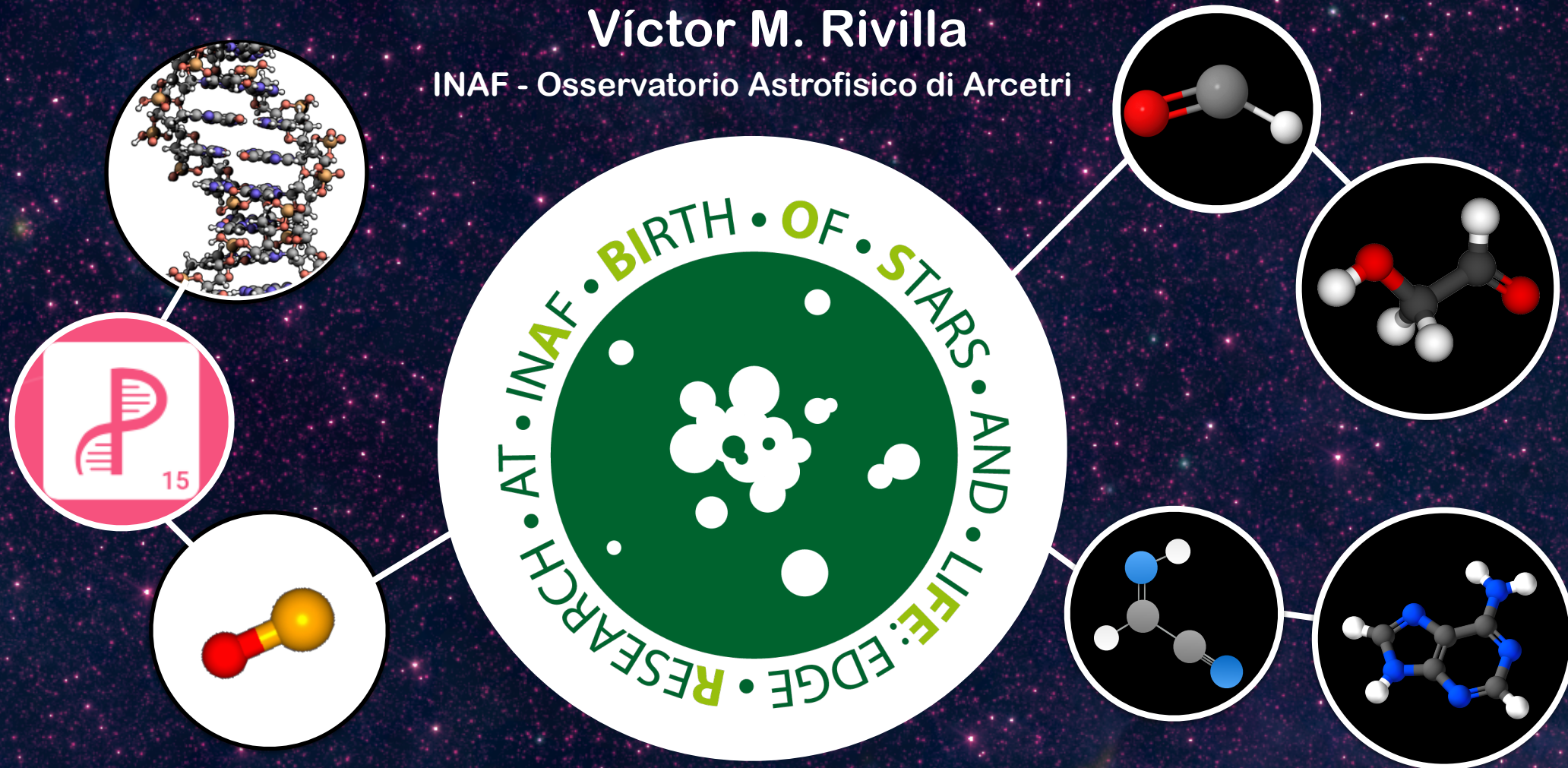


# BIOSFERA

Birth Of Stars and LiFE: Edge Research at INAF

Víctor M. Rivilla

INAF - Osservatorio Astrofisico di Arcetri



3rd ASTROFIT2 annual meeting  
Rome, Italy, October 16 2019

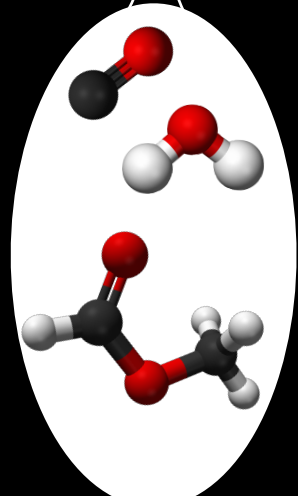
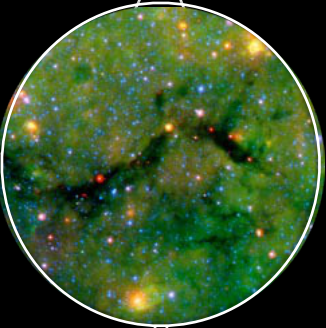
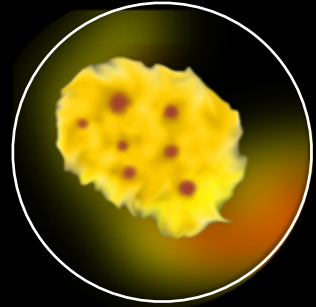


This project has received funding from the European Union's Horizon 2020 research and innovation programme under the Marie Skłodowska-Curie grant agreement No 664931.

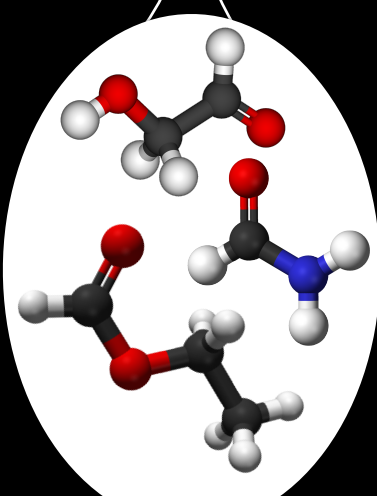
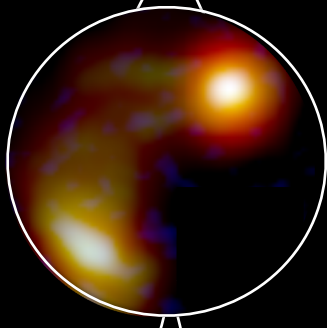
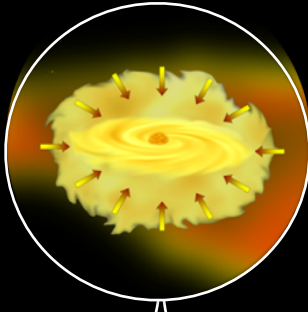


# From atoms & Simple molecules to ... LIFE

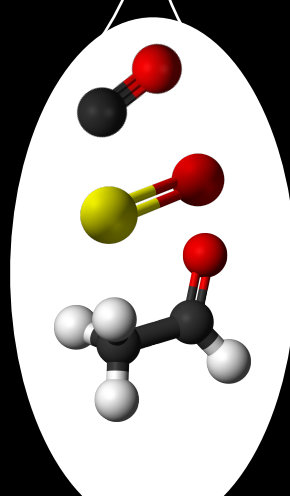
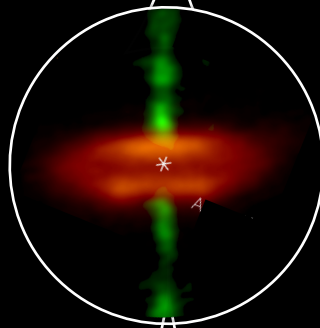
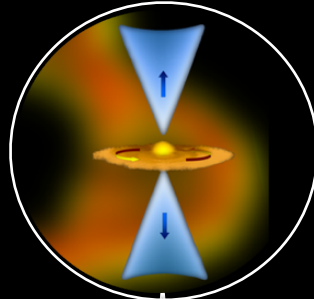
prestellar  
cores



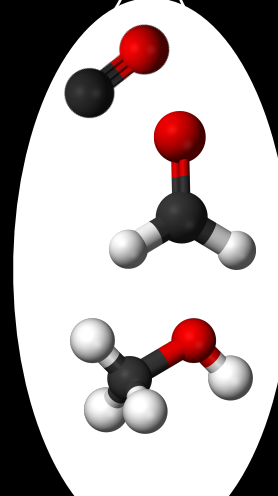
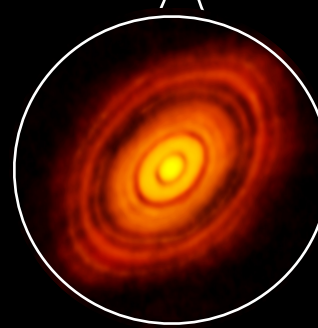
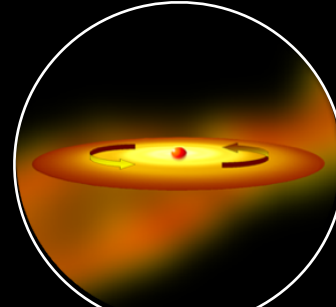
protostar  
envelope



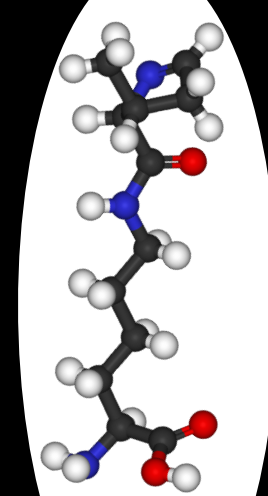
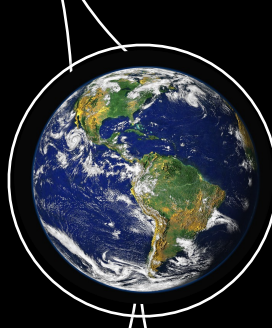
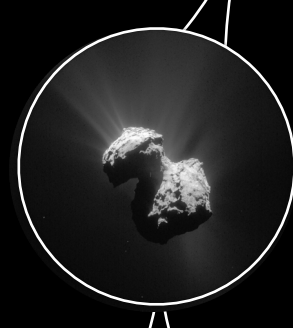
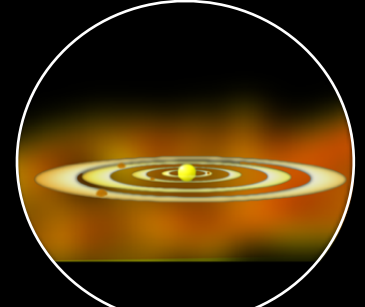
circumstellar  
disk and jet



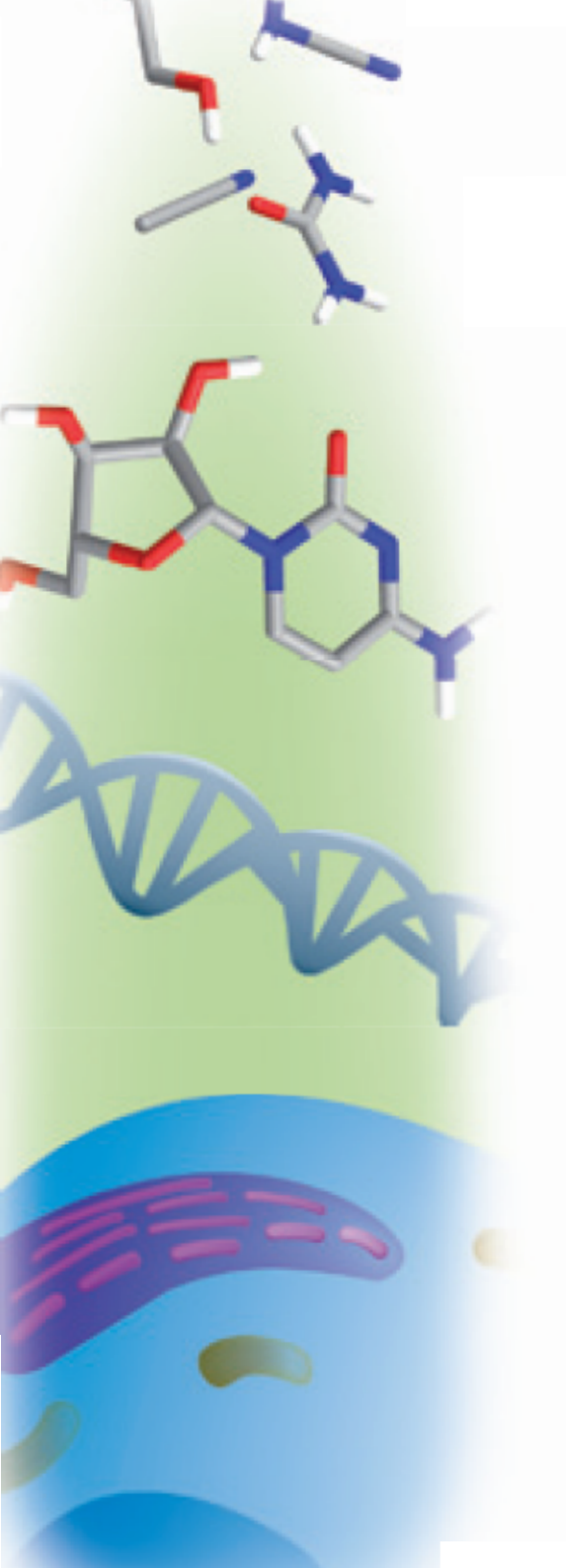
protoplanetary  
disk



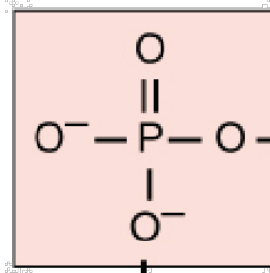
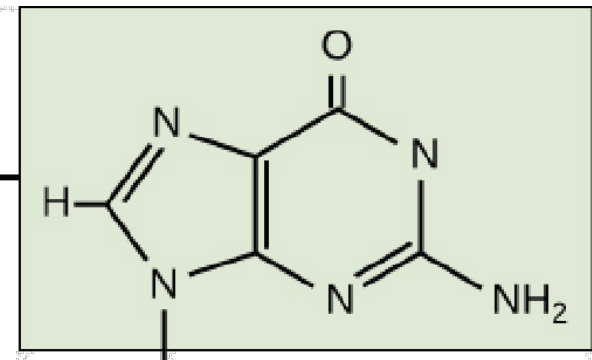
planets,  
comets & life



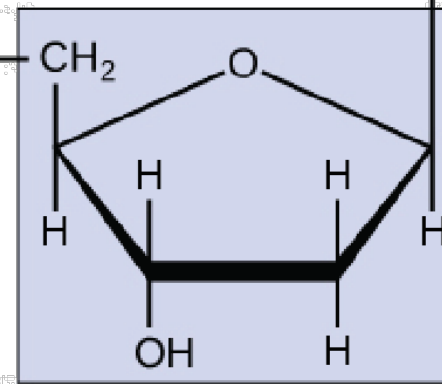




**nucleobases**



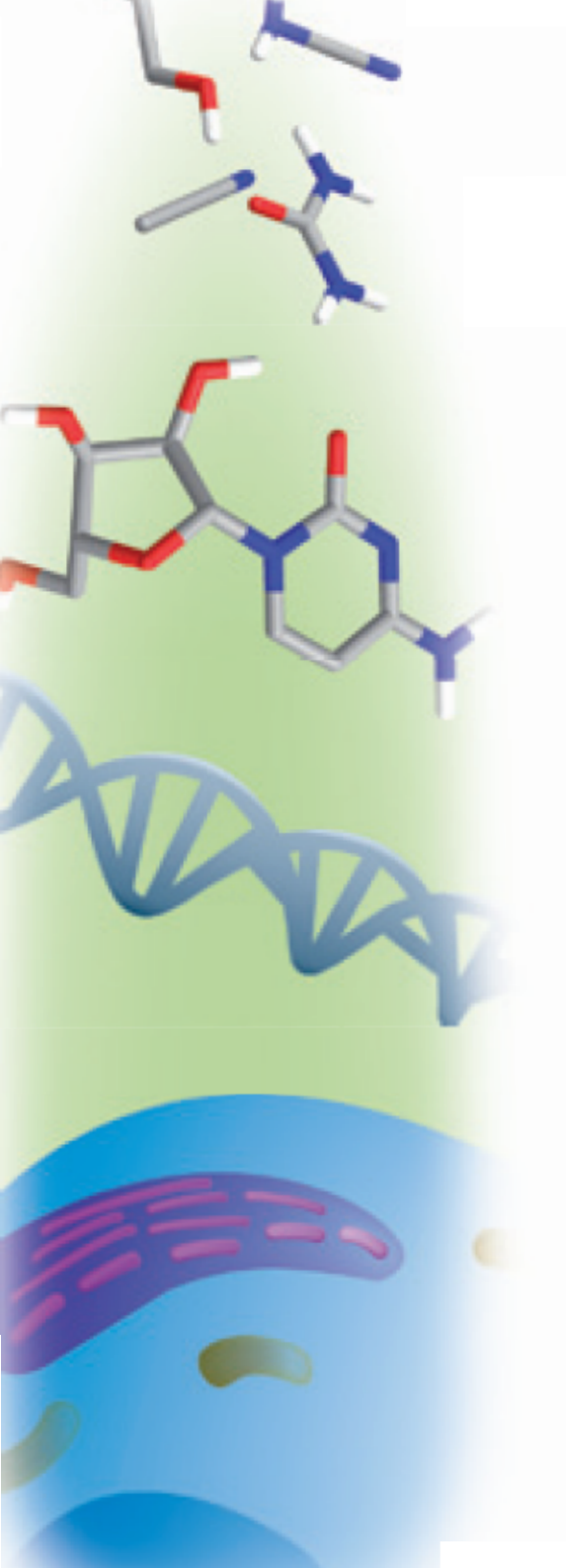
**phosphate**



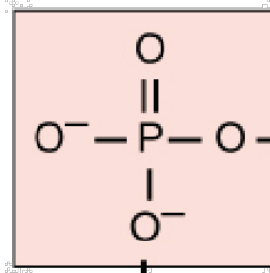
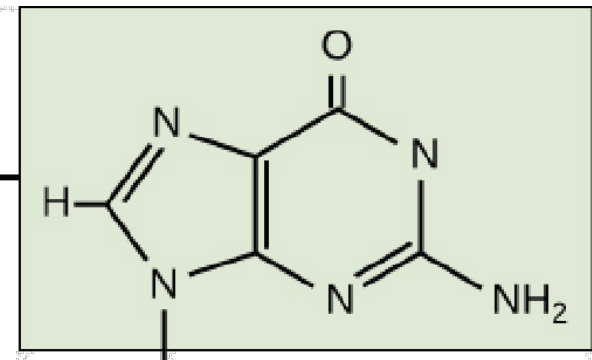
**sugars**

**Nucleotides:  
the building blocks of Life**

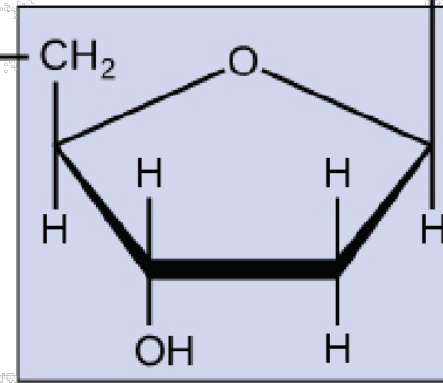




nucleobases



phosphate



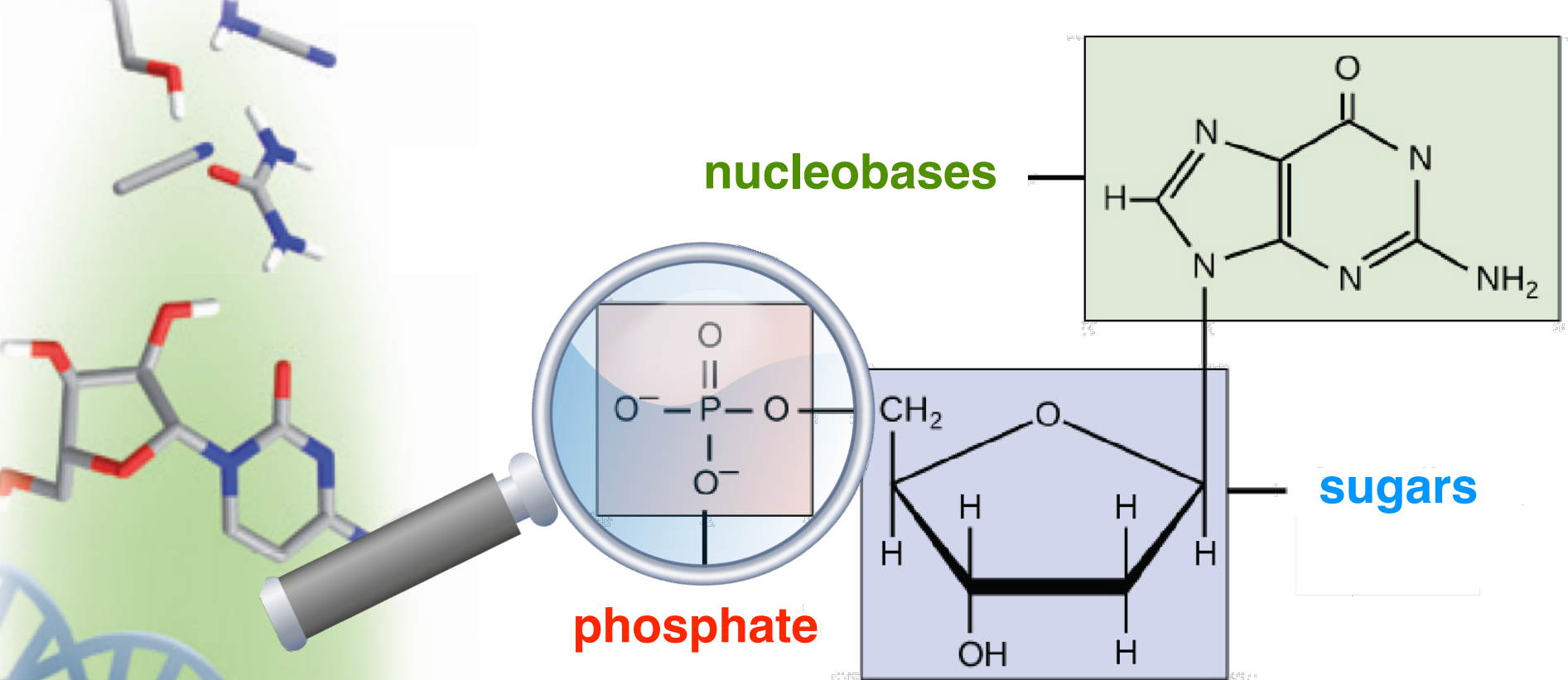
sugars

# BIOSFERA

Birth Of Stars and LiFE: Edge Research at INAF







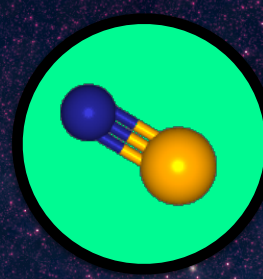
# BIOSFERA

Birth Of Stars and LiFE: Edge Research at INAF





# PN in the Galaxy



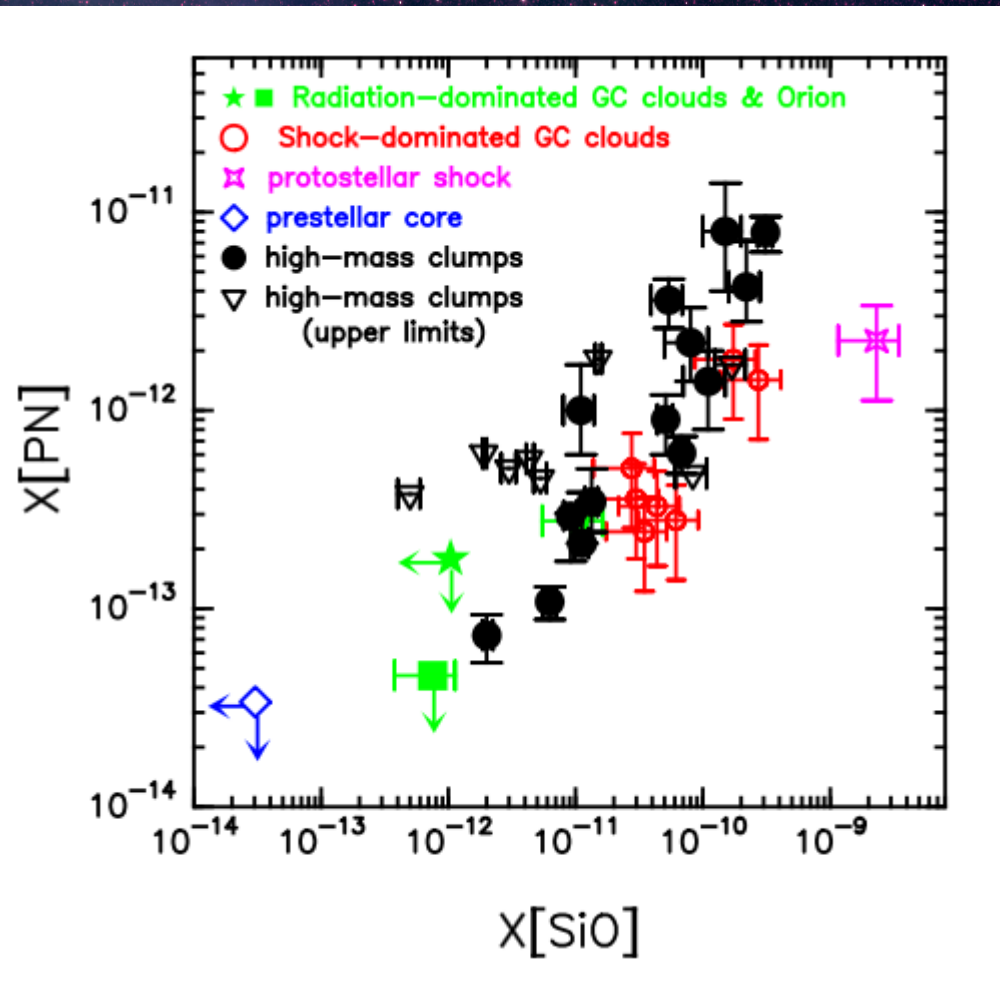
Fontani, Rivilla et al. (2016)

Rivilla et al. (2018)

Mininni, Fontani, Rivilla et al. (2018)



Survey of PN towards a large sample of massive star forming regions and Galactic Center clouds

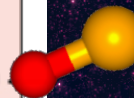
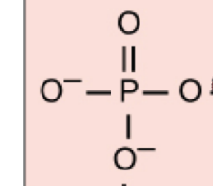






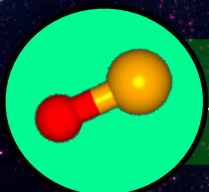
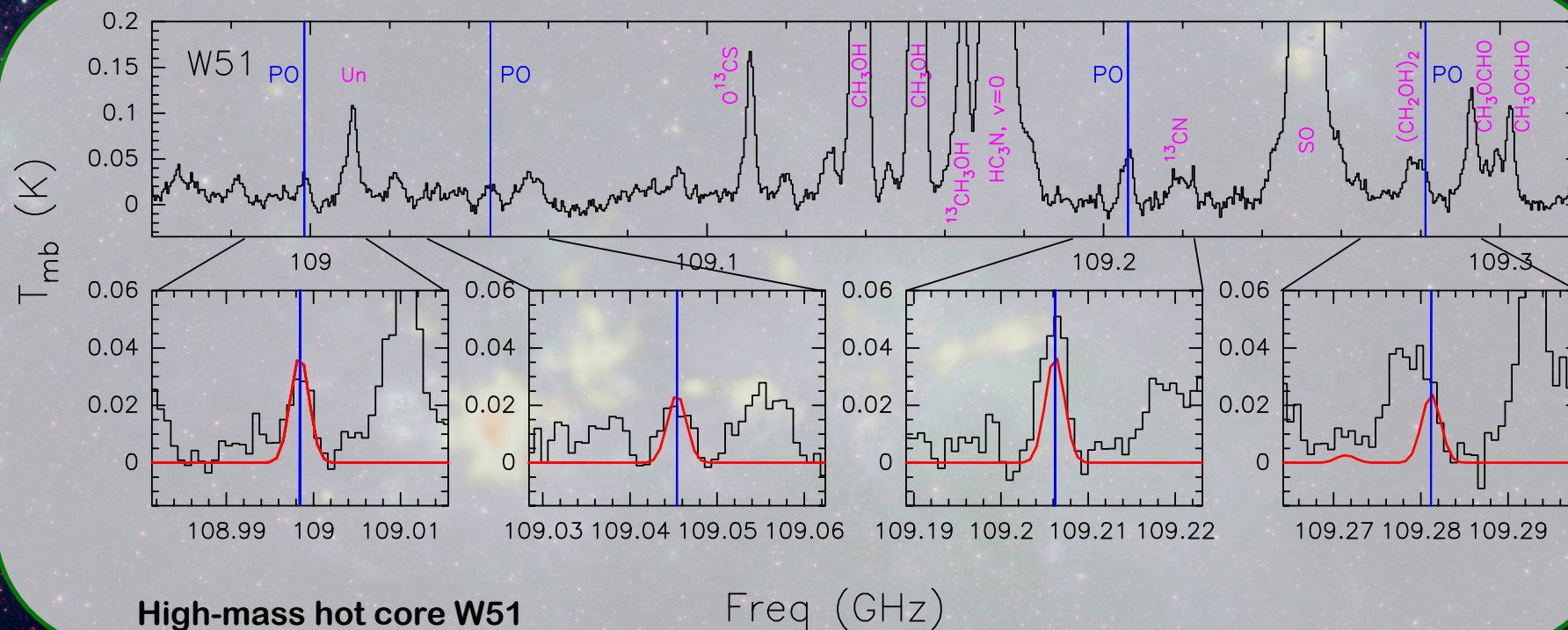
et al. (2016)

# First detections of PO in star-forming regions

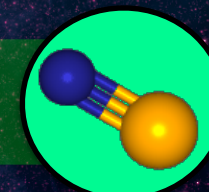


phosphate

Rivilla et al. 2016



PO is a factor 2-3 more abundant than PN !





# AFGL 5142 star-forming region



SO

Rivilla et al. (2019c)

## ALMA and ROSINA detections of phosphorus-bearing molecules: the interstellar thread between star-forming regions and comets

V. M. Rivilla<sup>1\*</sup>, M. N. Drozdovskaya<sup>2</sup>, K. Altwegg<sup>3</sup>, P. Caselli<sup>4</sup>, M. T. Beltrán<sup>1</sup>,  
F. Fontani<sup>1</sup>, F.F.S. van der Tak<sup>5,6</sup>, R. Cesaroni<sup>1</sup>, A. Vasyunin<sup>7,8</sup>, M. Rubin<sup>2</sup>, F. Lique<sup>9</sup>,  
S. Marinakis<sup>10,11</sup>, L. Testi<sup>1,12,13</sup>, and the ROSINA team<sup>14</sup>

*The full list of affiliations appears at the end of the paper in Appendix A.*

[-12, -4.1] km/s    $v_{\text{sys}}=3.8 \text{ km/s}$    [-3.5, 6.0] km/s

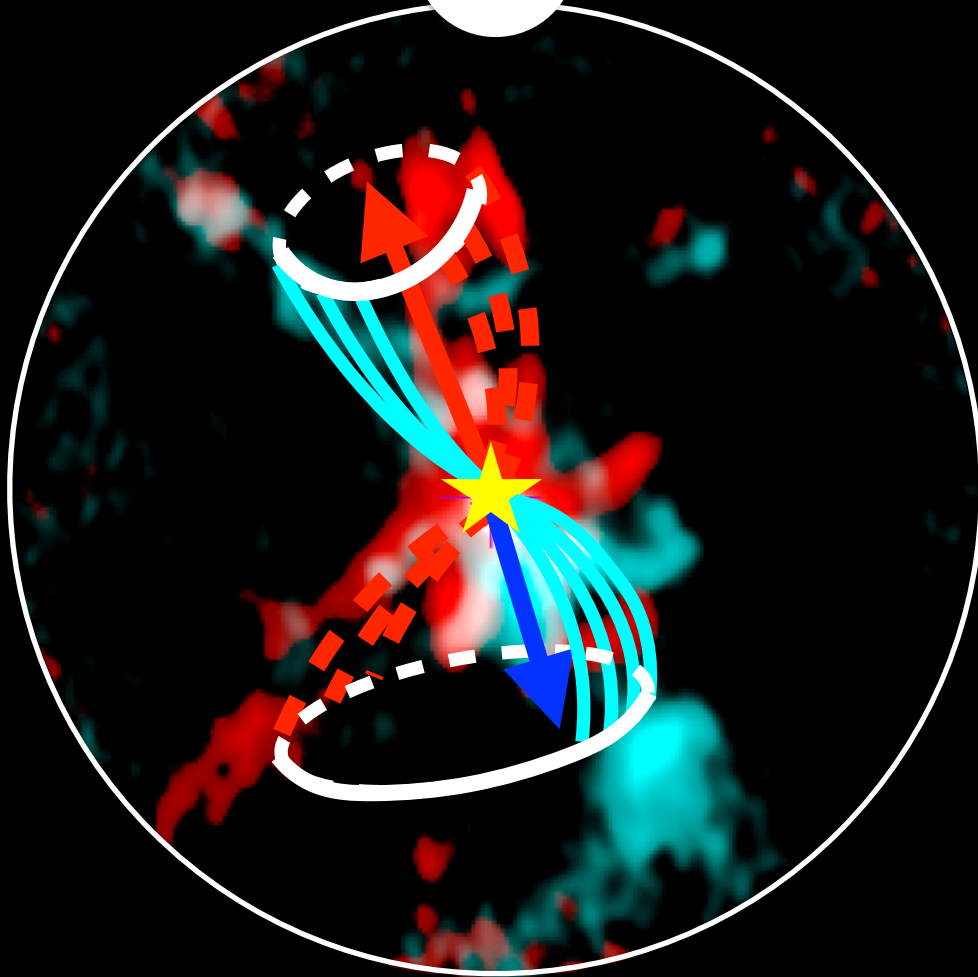


# AFGL 5142 star-forming region



SO

Rivilla et al. (2019c)



$[-12, -4.1] \text{ km/s}$     $v_{\text{sys}}=3.8 \text{ km/s}$     $[-3.5, 6.0] \text{ km/s}$

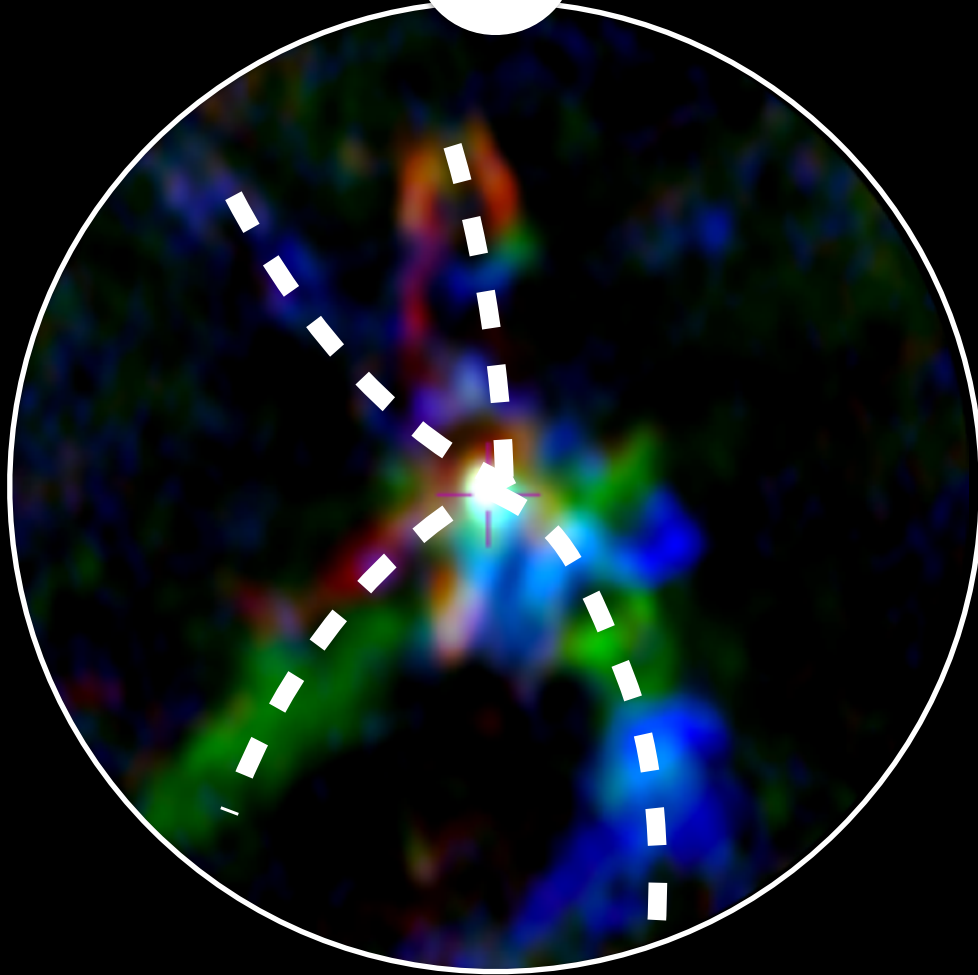


# AFGL 5142 star-forming region



SO

Rivilla et al. (2019c)



$[-6.5, -4.5]$  km/s    $[-4.5, -2.0]$  km/s    $[-2.0, 0.0]$  km/s



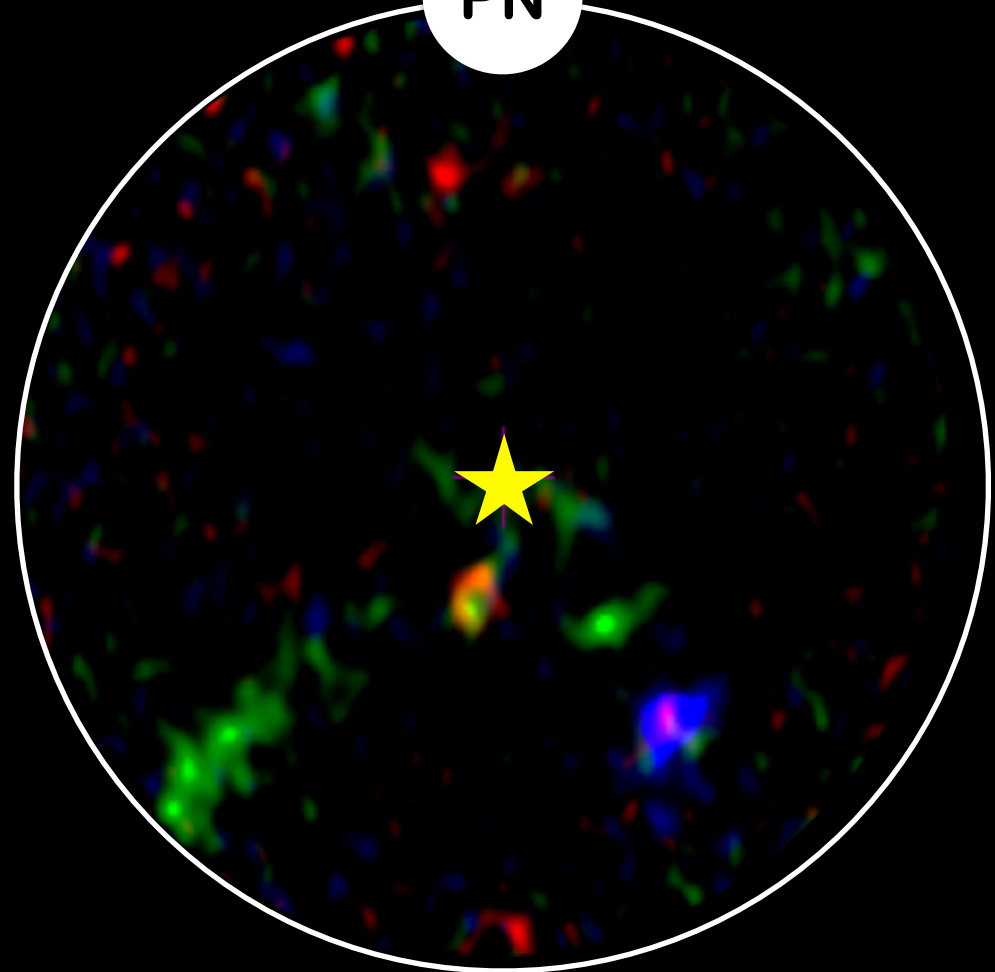
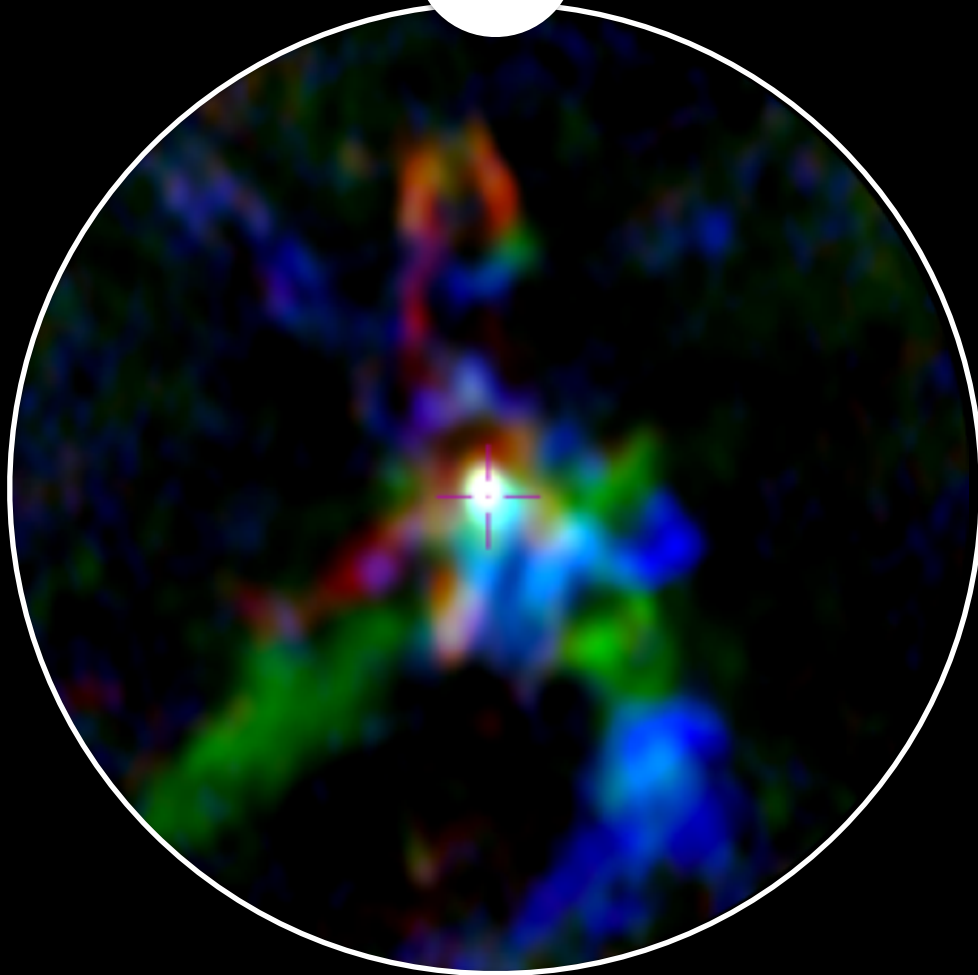
# AFGL 5142 star-forming region



SO

Rivilla et al. (2019c)

PN



[-6.5, -4.5] km/s   [-4.5, -2.0] km/s   [-2.0, 0.0] km/s   [-6.5, -4.5] km/s   [-4.5, -2.0] km/s   [-2.0, 0.0] km/s



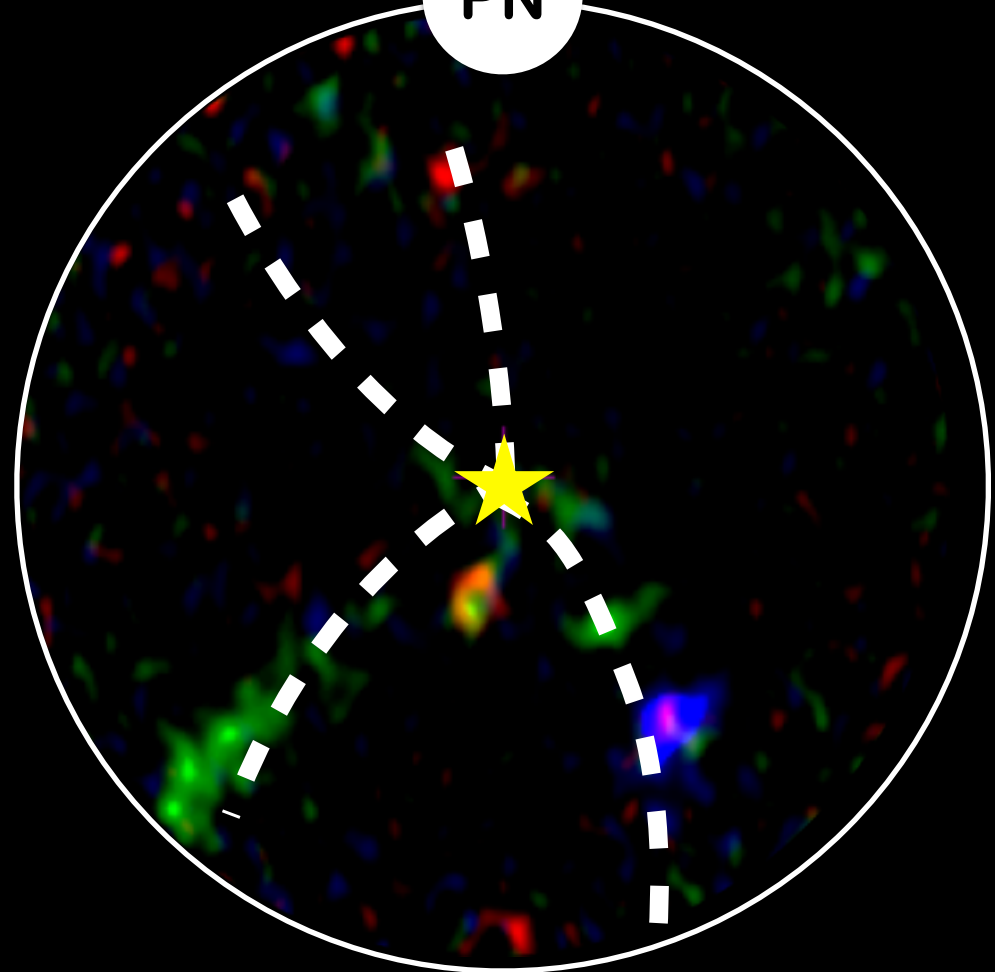
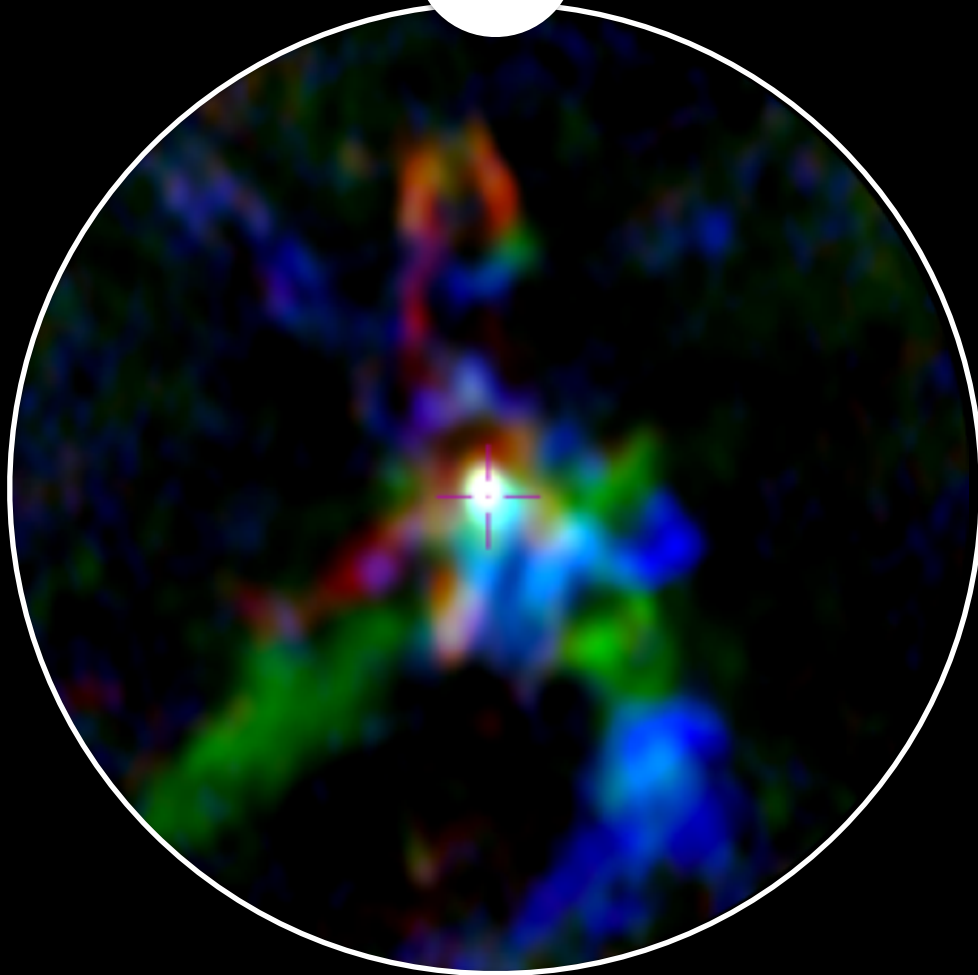
# AFGL 5142 star-forming region



SO

Rivilla et al. (2019c)

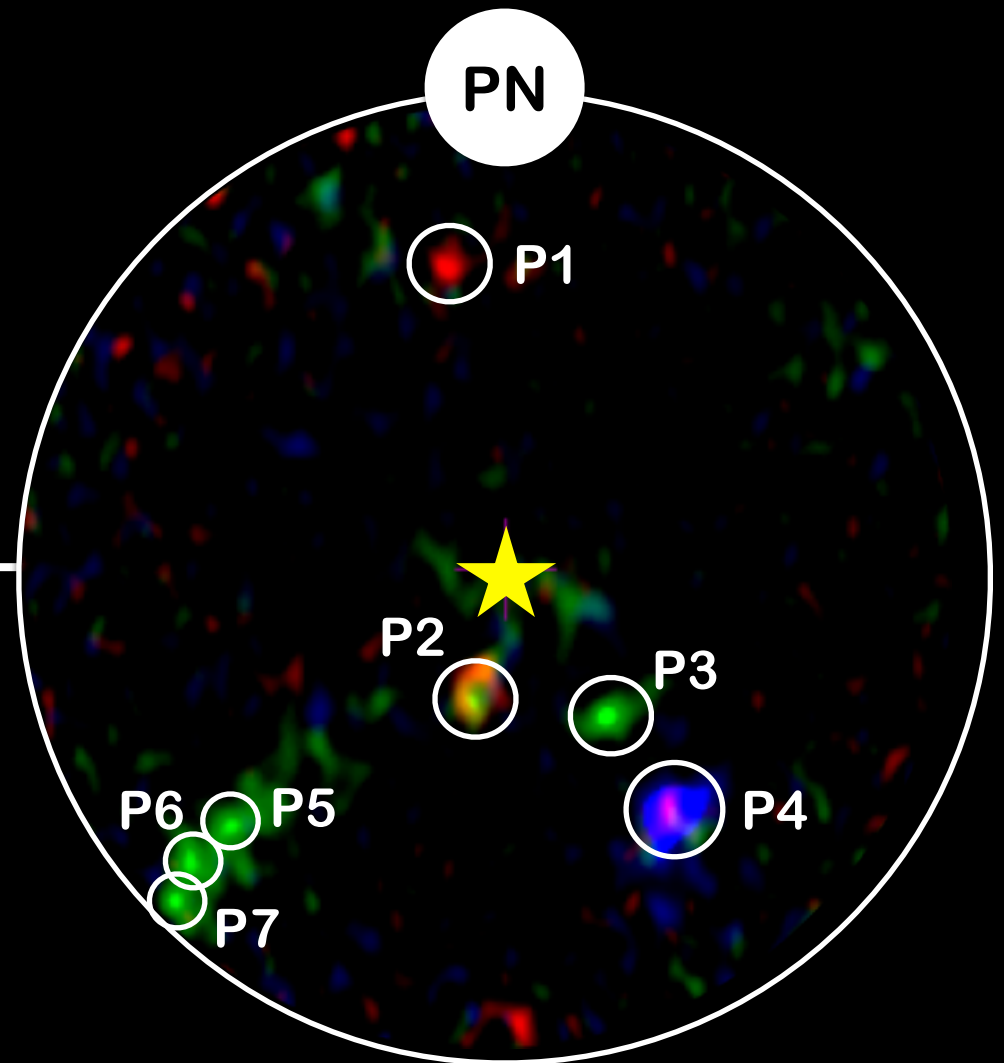
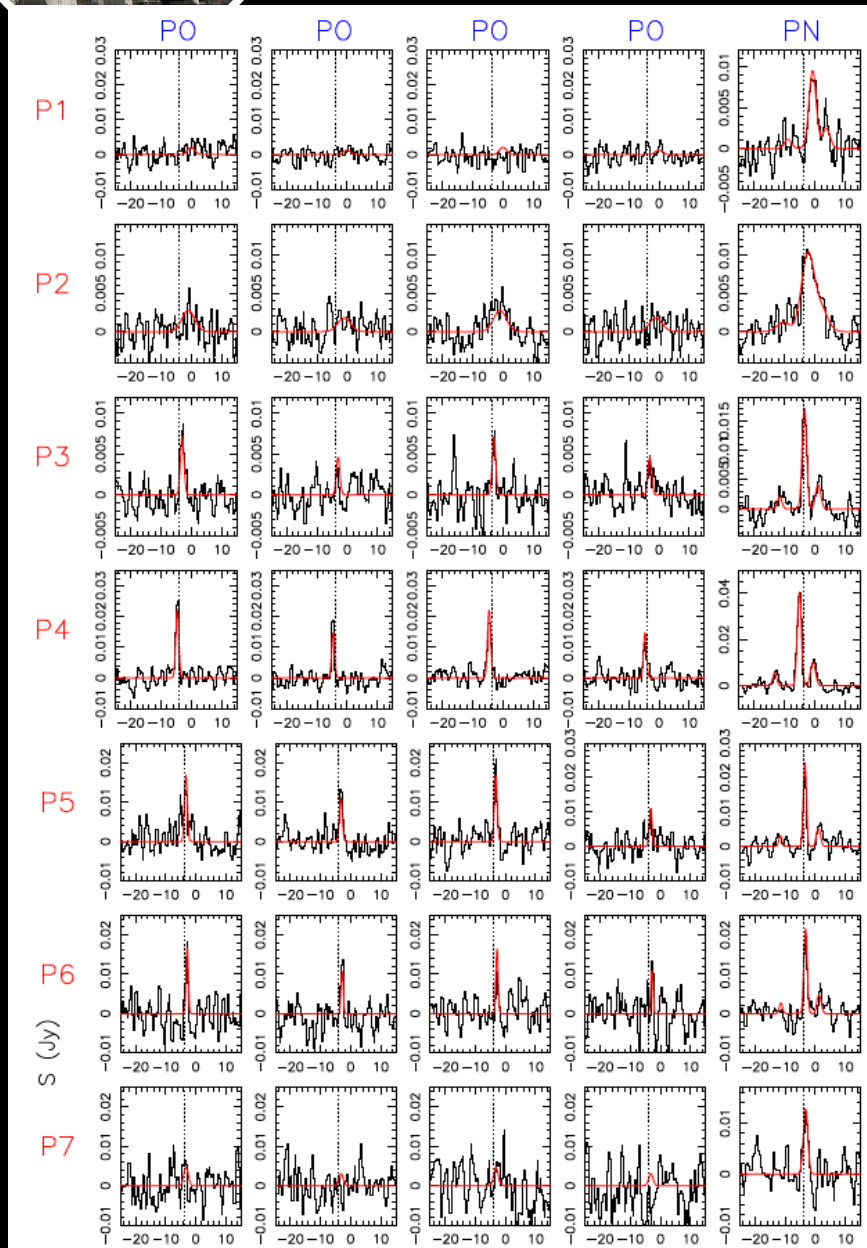
PN



[-6.5, -4.5] km/s [-4.5, -2.0] km/s [-2.0, 0.0] km/s [-6.5, -4.5] km/s [-4.5, -2.0] km/s [-2.0, 0.0] km/s



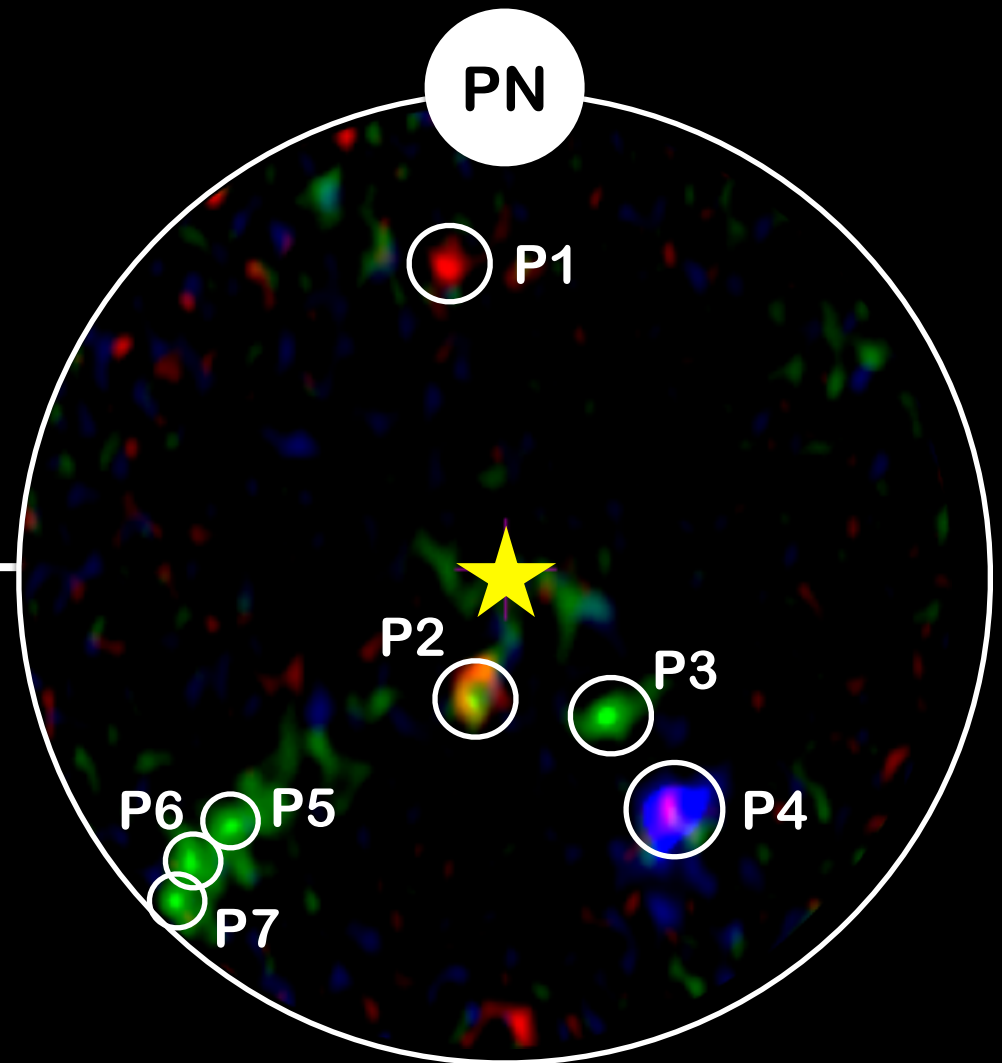
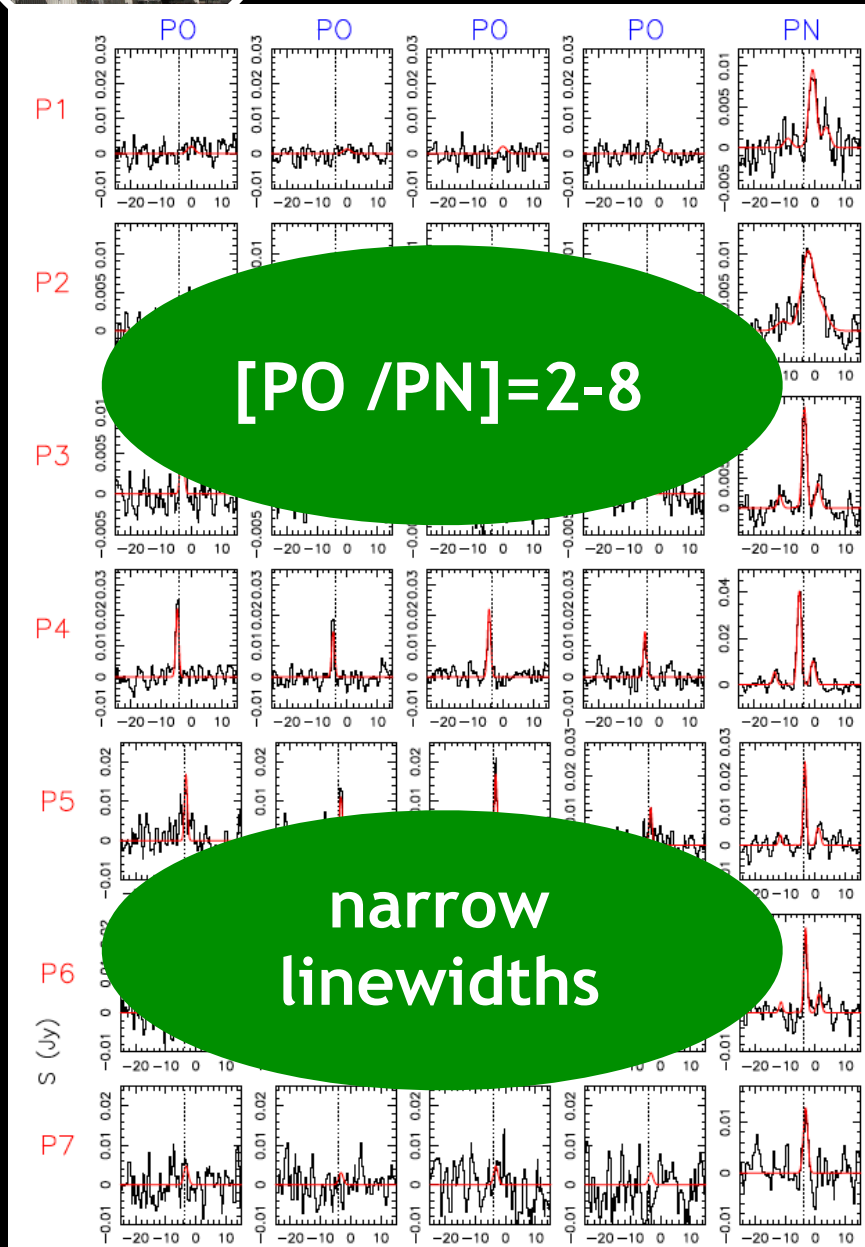
# AFGL 5142 star-forming region



[ -6.5, -4.5 ] km/s   [ -4.5, -2.0 ] km/s   [ -2.0, 0.0 ] km/s



# AFGL 5142 star-forming region

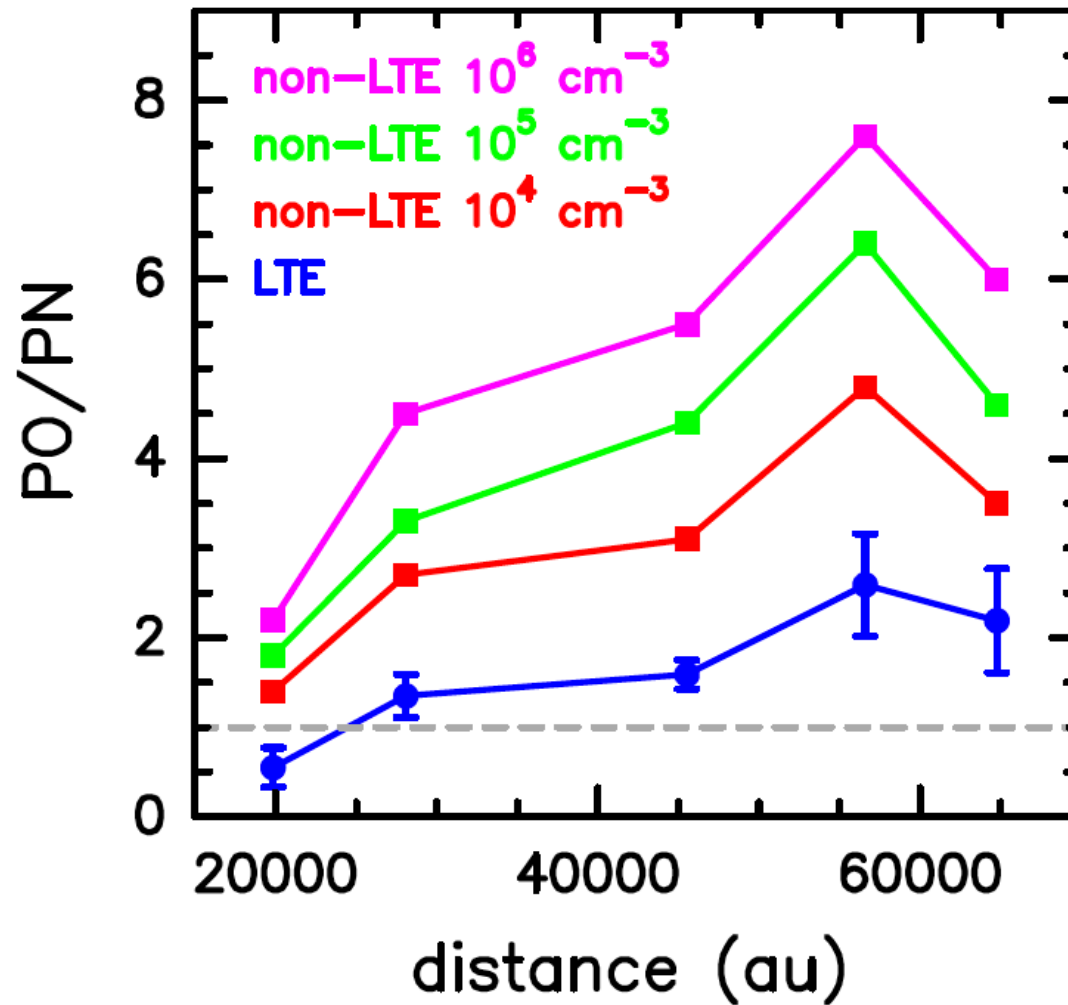


$[-6.5, -4.5]$  km/s    $[-4.5, -2.0]$  km/s    $[-2.0, 0.0]$  km/s



# PO/PN ratio

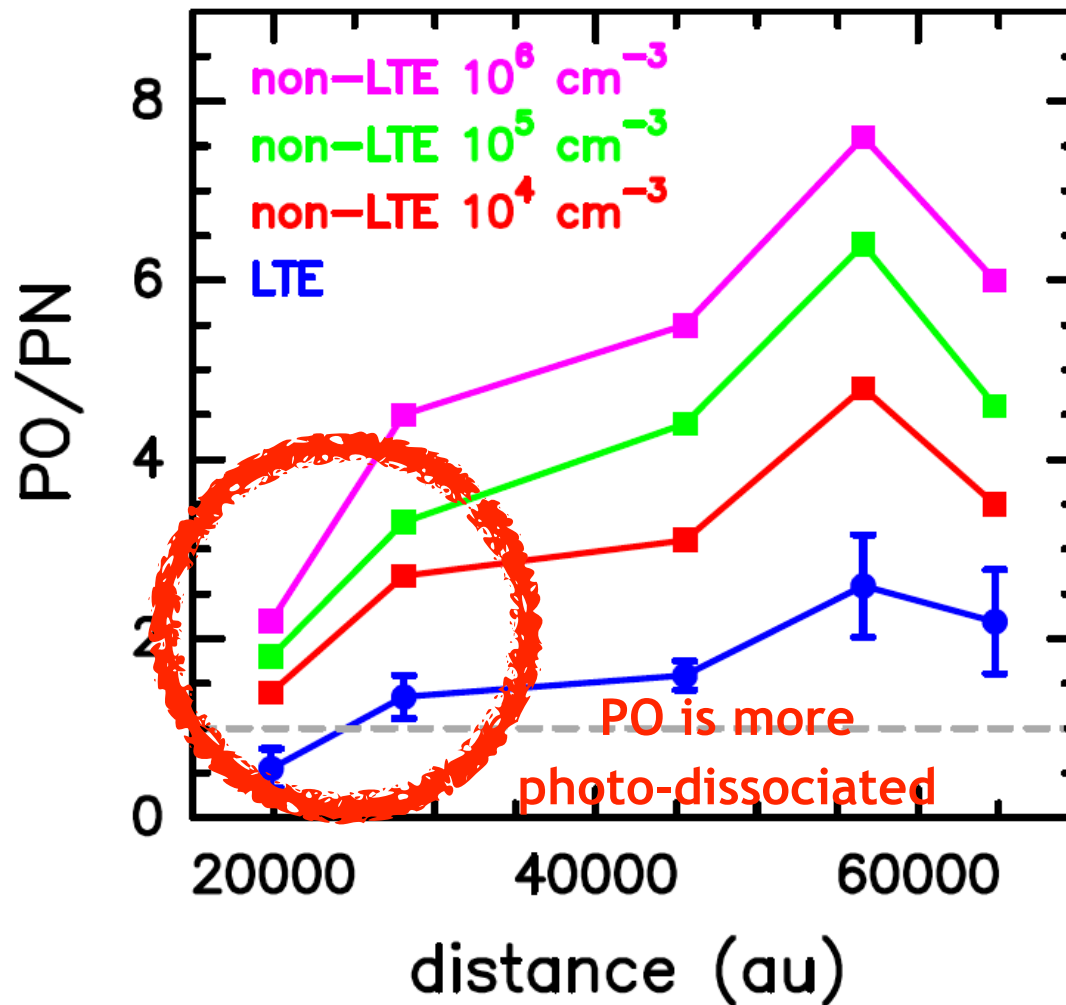
Rivilla et al. (2019c)





# PO/PN ratio

Rivilla et al. (2019c)





# Formation scenario

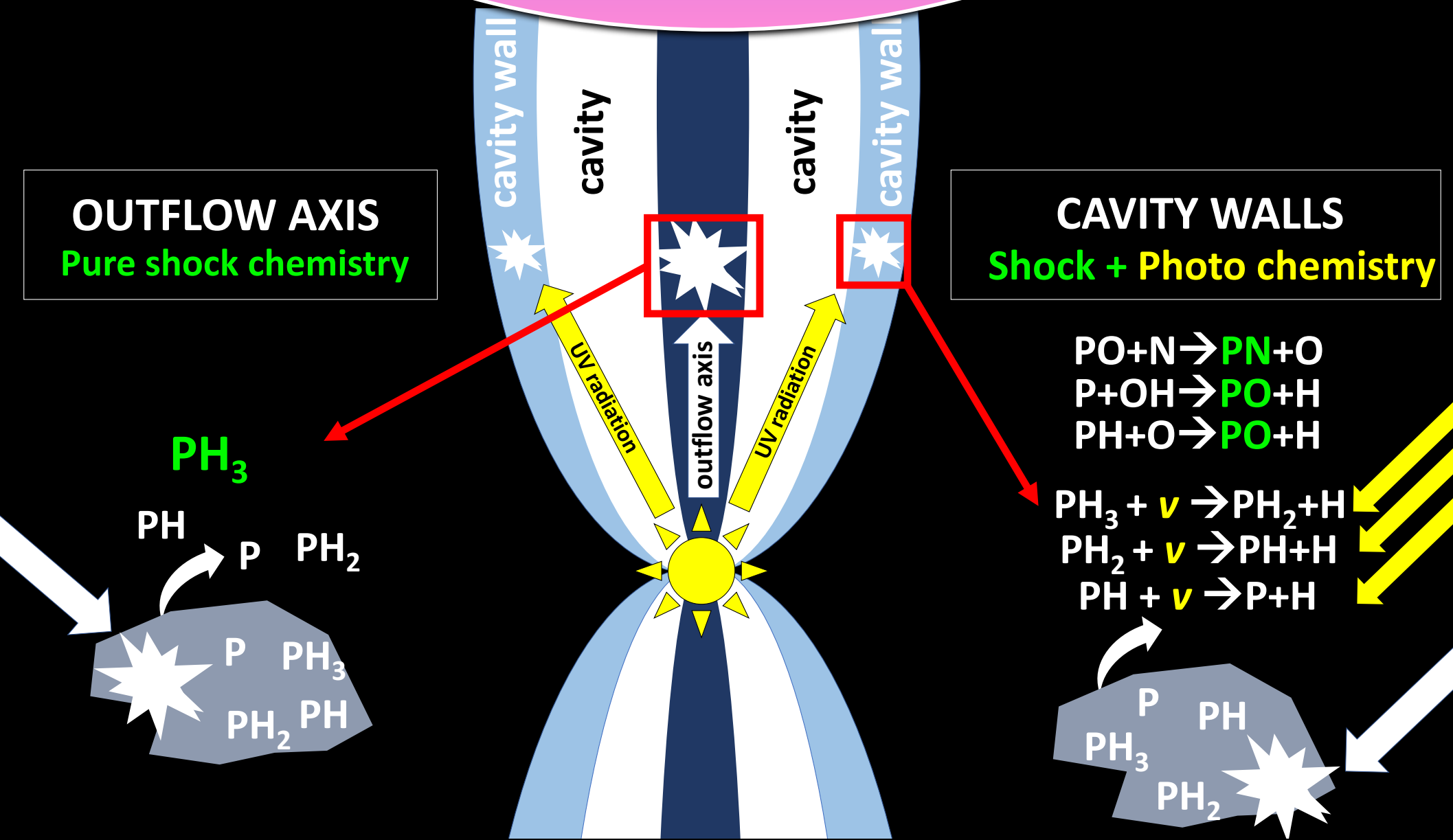
Shocks + photochemistry

OUTFLOW AXIS

Pure shock chemistry

CAVITY WALLS

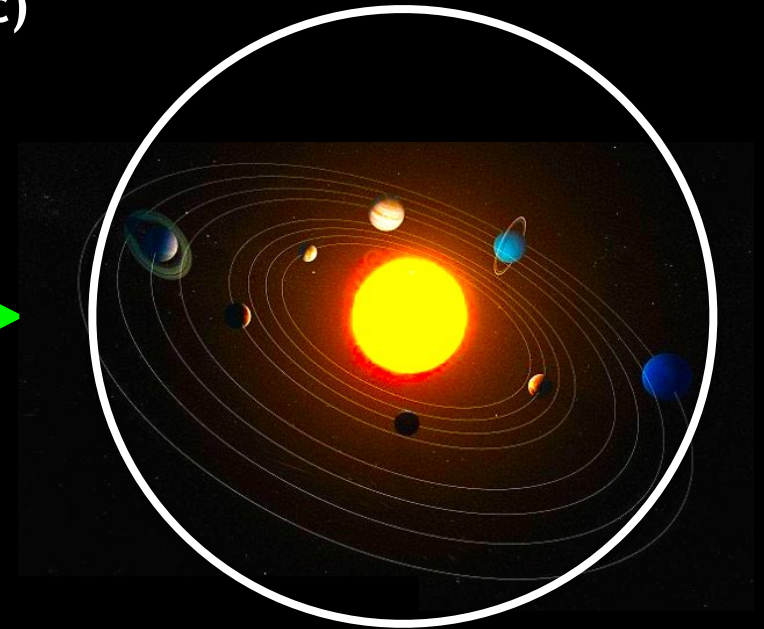
Shock + Photo chemistry





# The Phosphorus connection between protostars and comets

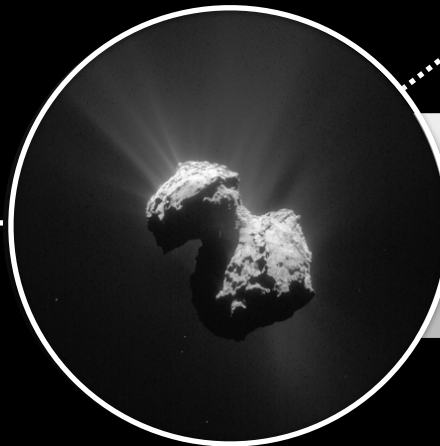
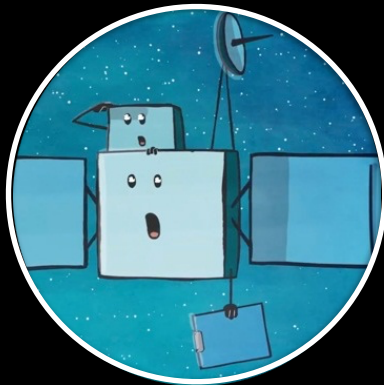
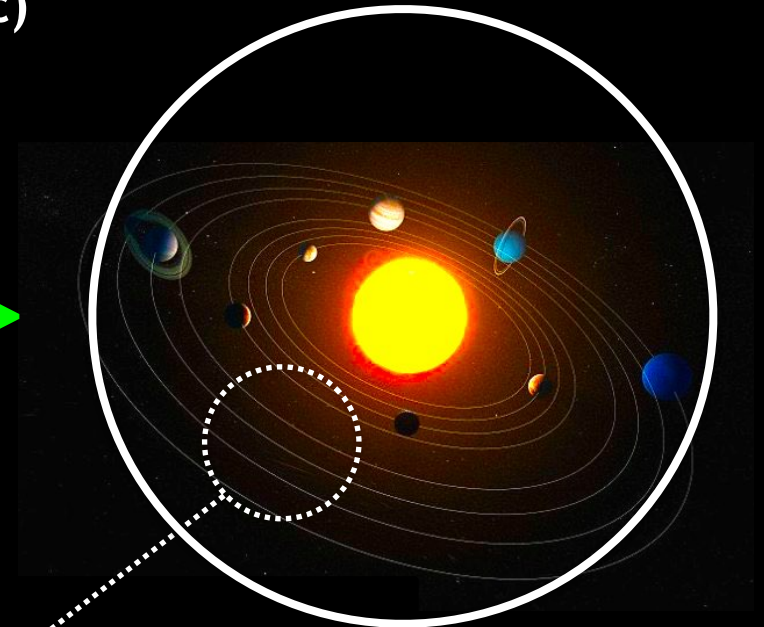
Rivilla et al. (2019c)





# The Phosphorus connection between protostars and comets

Rivilla et al. (2019c)

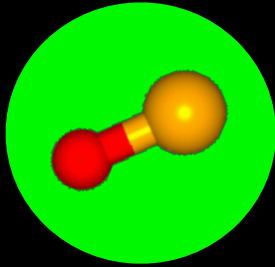
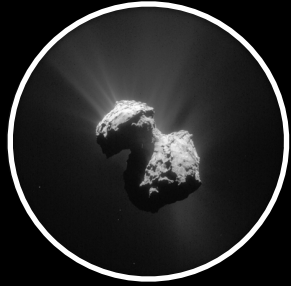


Pristine material from the  
early Solar System.

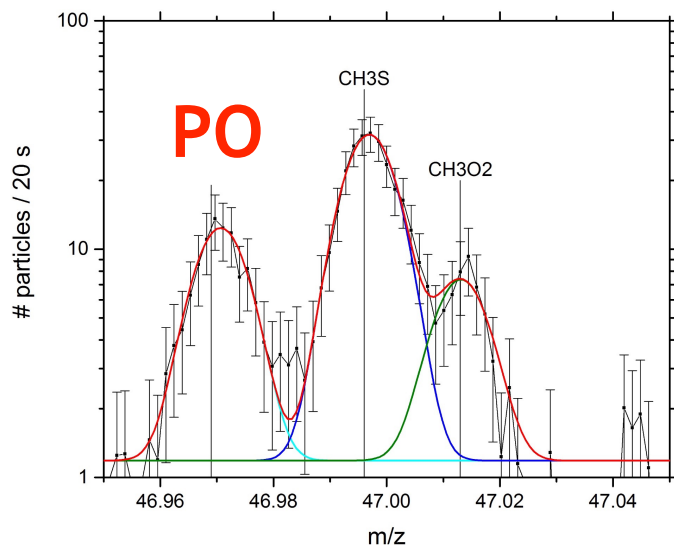
67P Churyumov-Gerasimenko comet



# Phosphorus in 67P



Rivilla et al. (2019c)



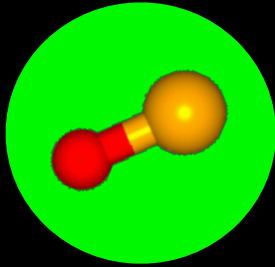
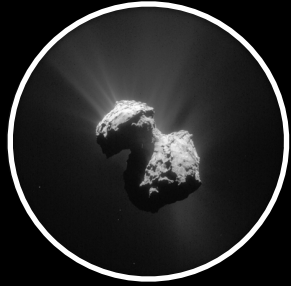
PO is present in the comet.

- Upper limits for PN, PH<sub>3</sub> and CP.

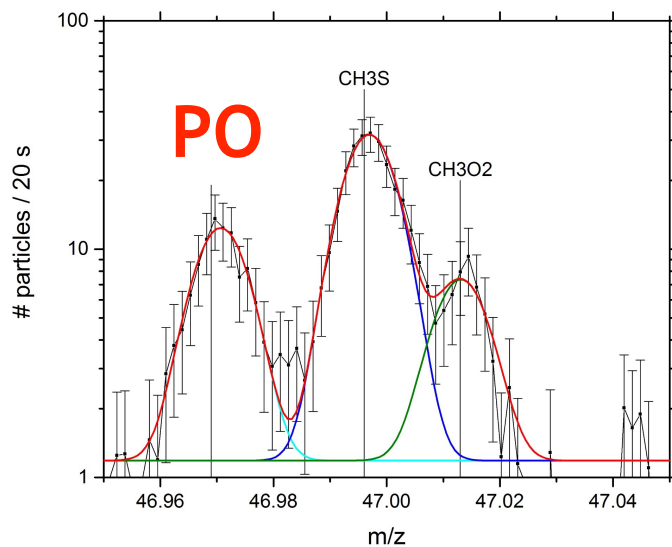
$[PO/PN] > 10$



# Phosphorus in 67P



Rivilla et al. (2019c)



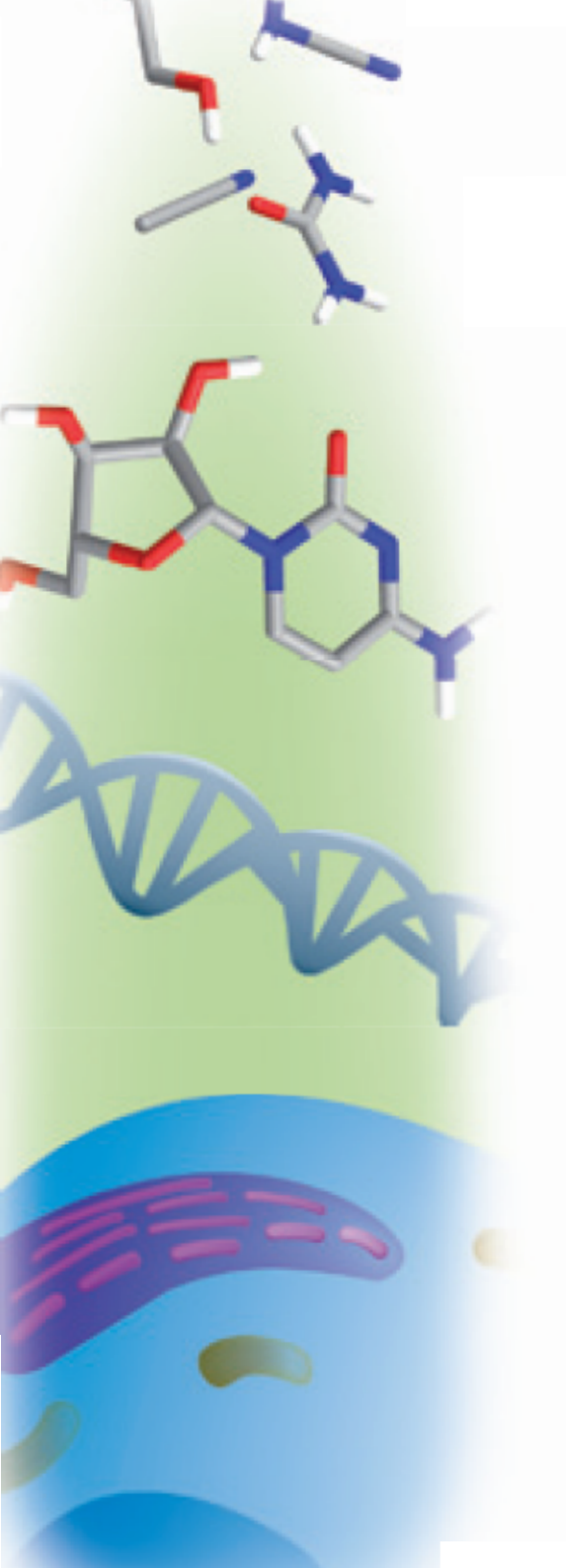
PO is present in the comet.

- Upper limits for PN, PH<sub>3</sub> and CP.

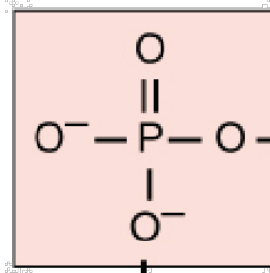
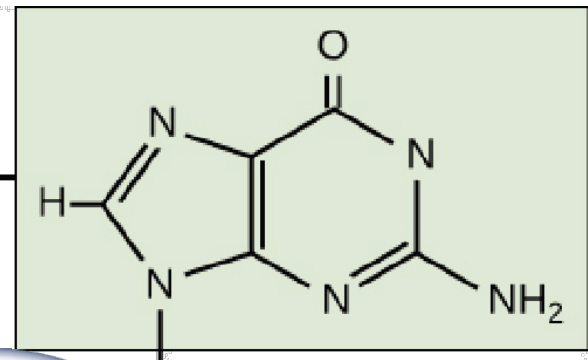
$[PO/PN] > 10$

PO is more abundant than PN both in star-forming regions and the comet.

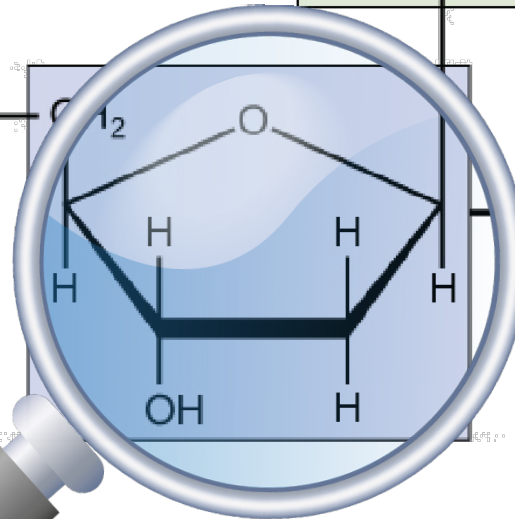




nucleobases



phosphate



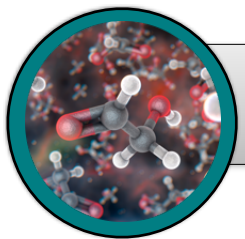
sugars

# BIOSFERA

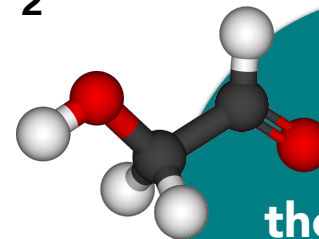
Birth of Stars and LiFE: Edge Research at INAF







# The formation of complex organic molecules: a complex problem

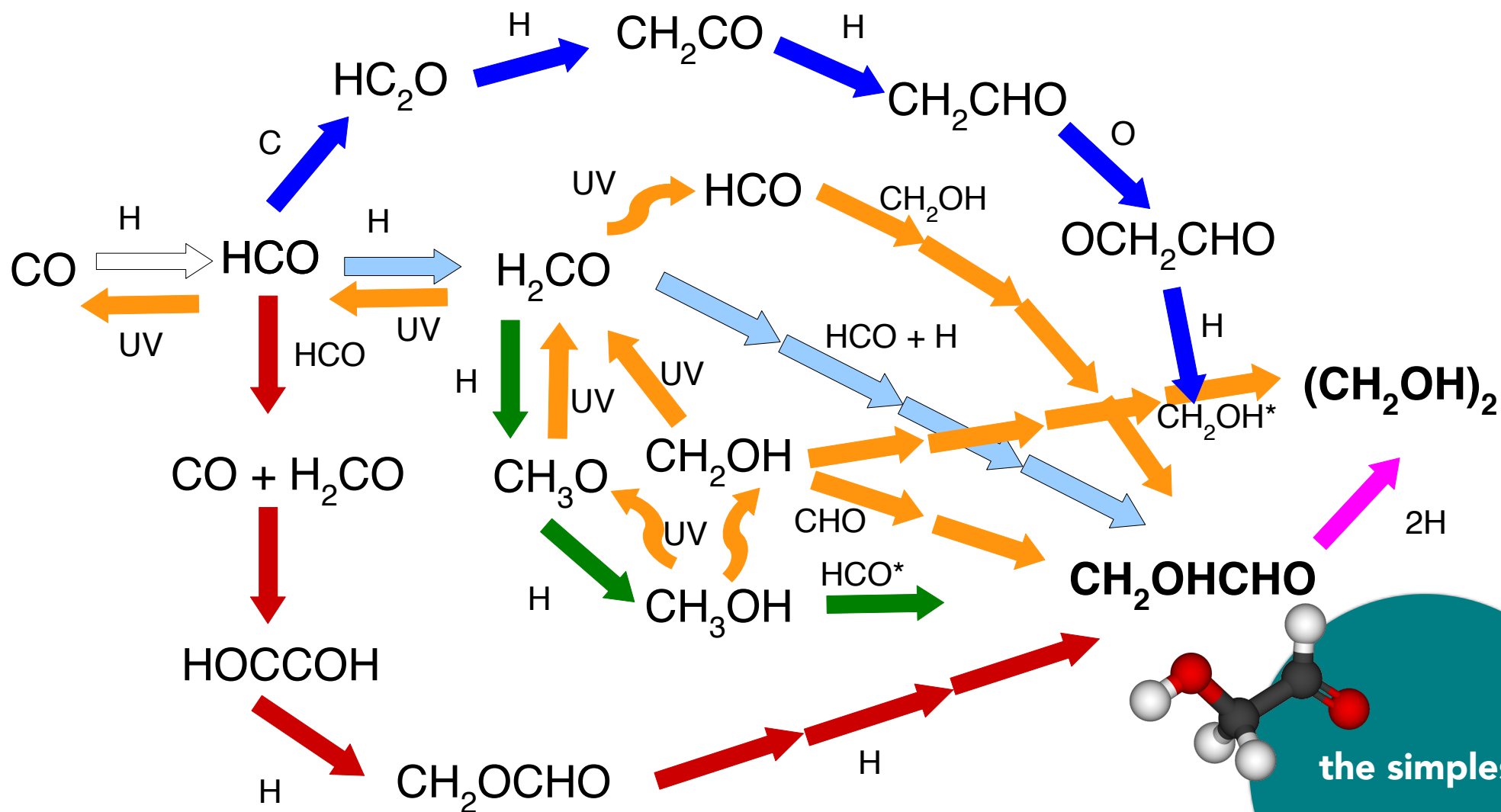


the simplest  
sugar



# The formation of complex organic molecules: a complex problem

## GAME OF REACTIONS



— Woods+13, Fedoseev+15

— Beltrán+09, Woods+12

— Charnley+1997, Sorrell+01

— Oberg+09

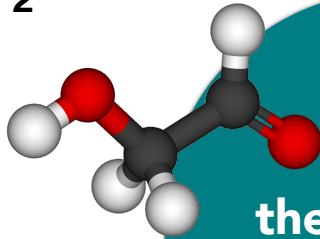
— Fedoseev+15

the simplest  
sugar





# GAME OF REACTIONS



**the simplest  
sugar**

Beltrán+09, Woods+12

— Oberg+09

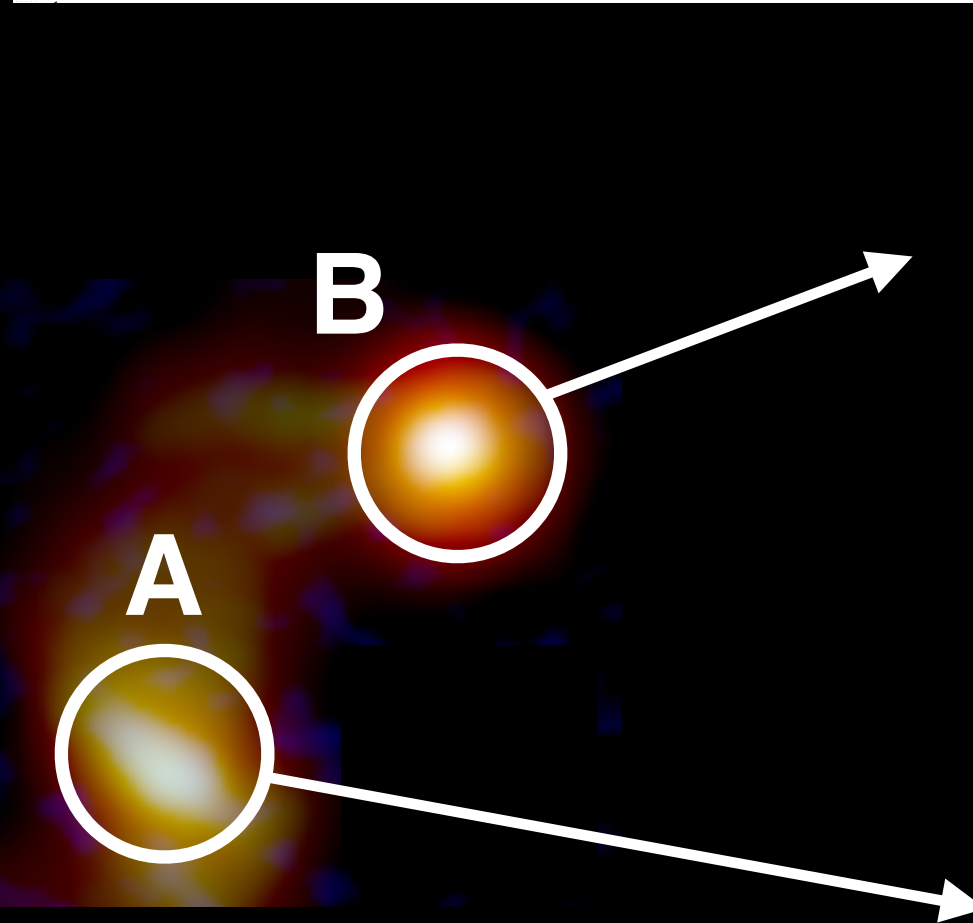
 Fedoseev+15



# First ALMA maps of HCO, an important precursor of complex organic molecules, towards IRAS 16293–2422

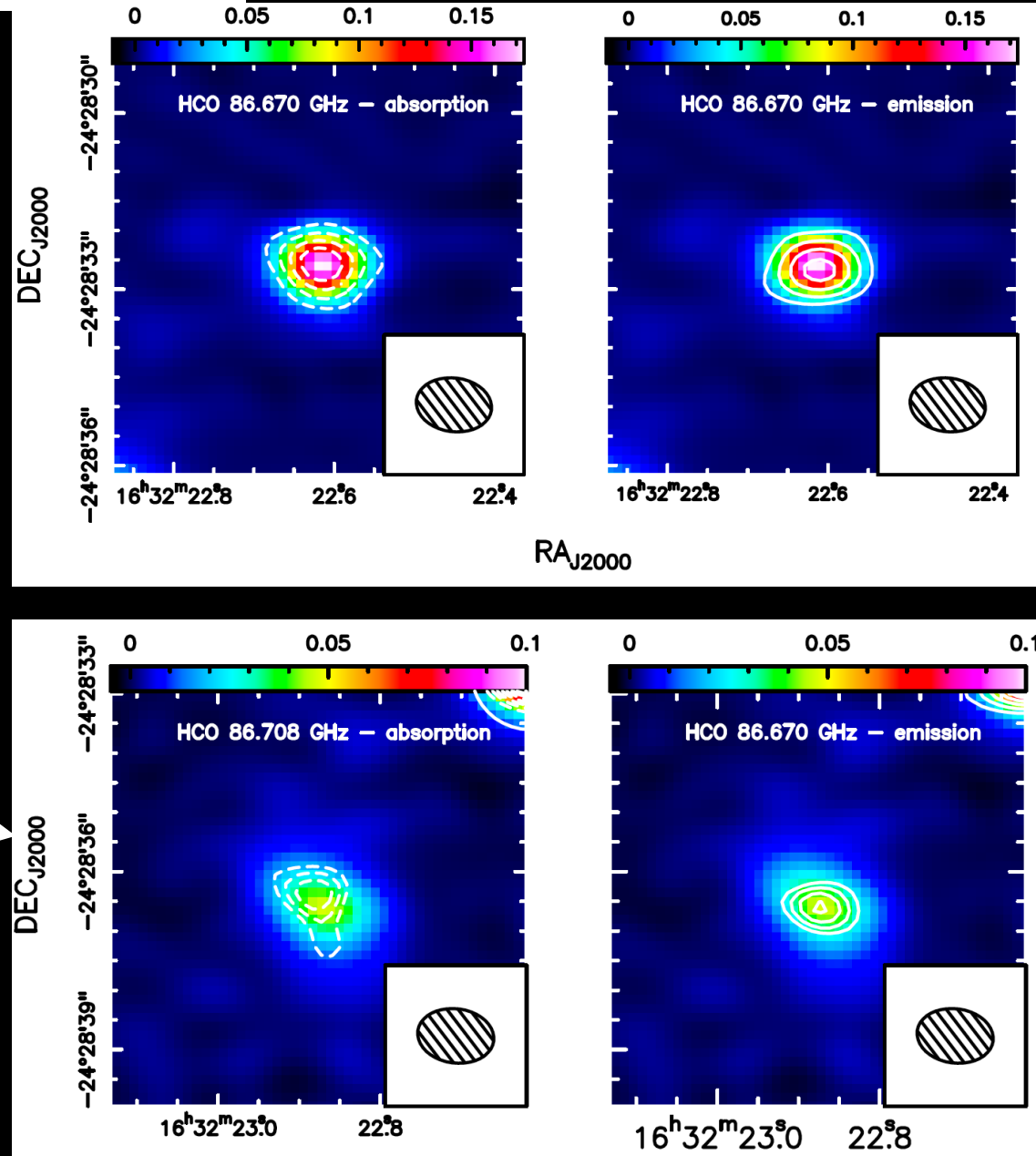
V. M. Rivilla<sup>1</sup>★, M. T. Beltrán<sup>1</sup>, A. Vasyunin<sup>2,3,4</sup>, P. Caselli<sup>2</sup>, S. Viti<sup>5</sup>,  
F. Fontani<sup>1</sup>, and R. Cesaroni<sup>1</sup>

Rivilla et al., 2019a

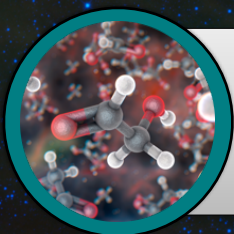


**ALMA MAPS**

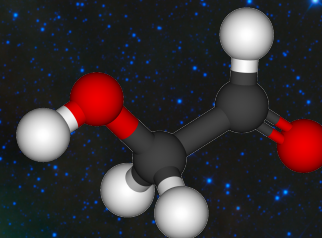
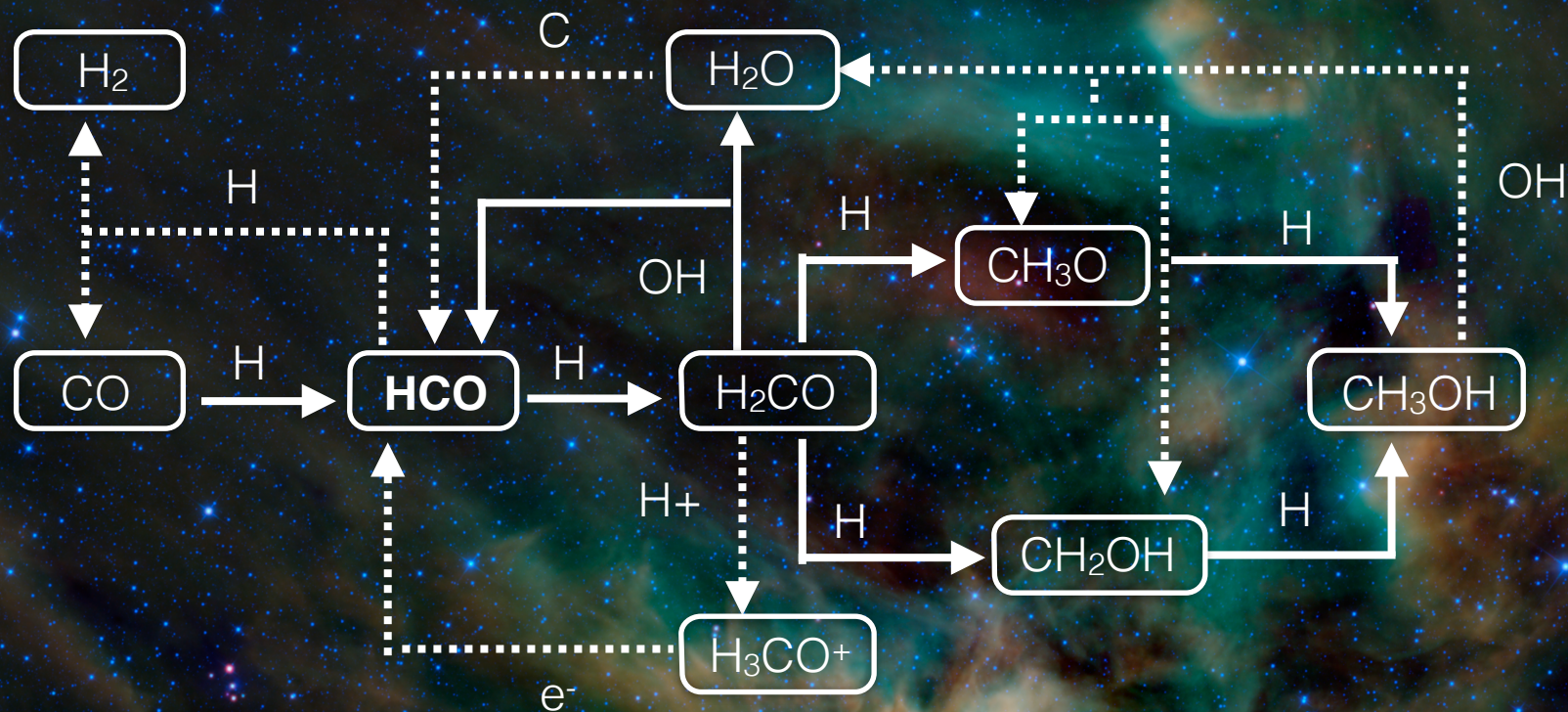
1"-1.5" resolution  
150 - 225 au





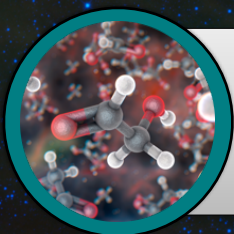


# The formation of glycolaldehyde $\text{CH}_2\text{OHCHO}$

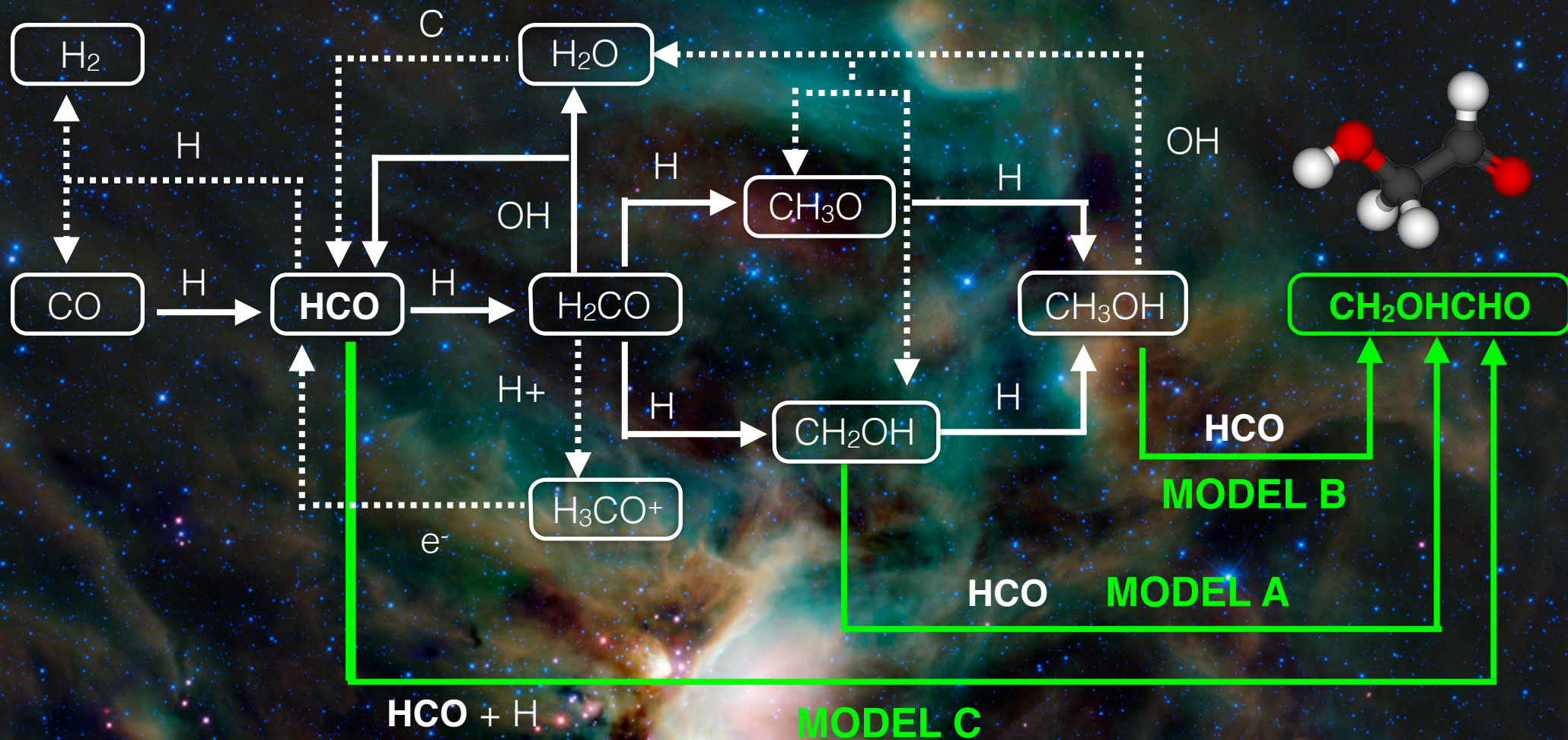


**$\text{CH}_2\text{OHCHO}$**

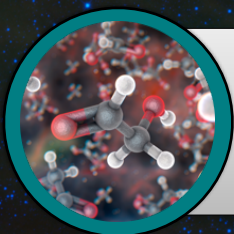




# The formation of glycolaldehyde $\text{CH}_2\text{OHCHO}$

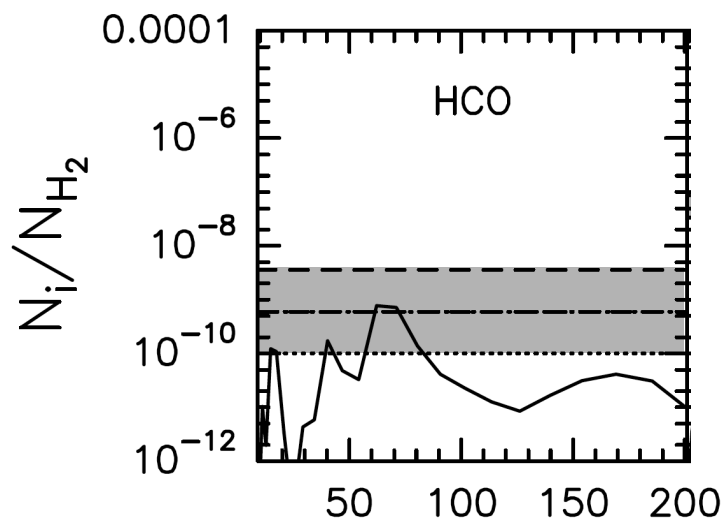




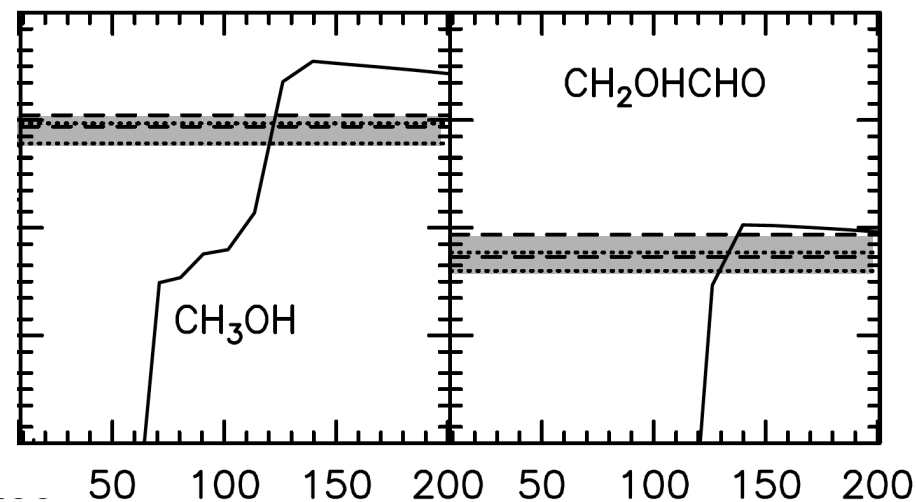
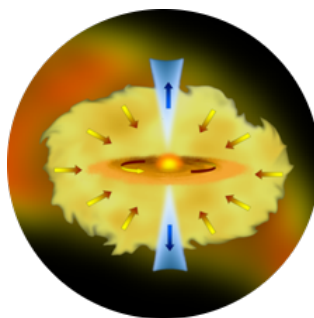


## The formation of glycolaldehyde $\text{CH}_2\text{OHCHO}$

- The model that fits better the observations is **MODEL A**:

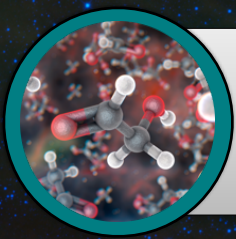


**MODEL A**



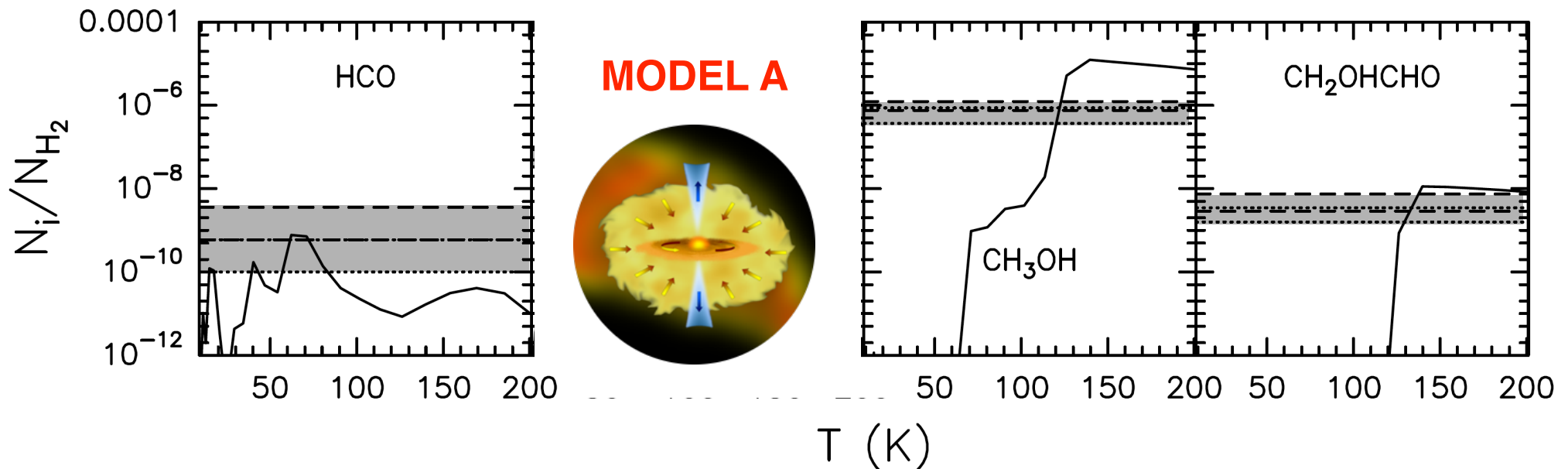
$T$  (K)





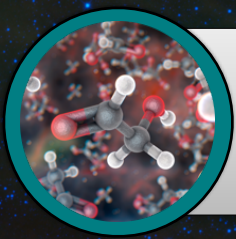
## The formation of glycolaldehyde $\text{CH}_2\text{OHCHO}$

- The model that fits better the observations is **MODEL A**:



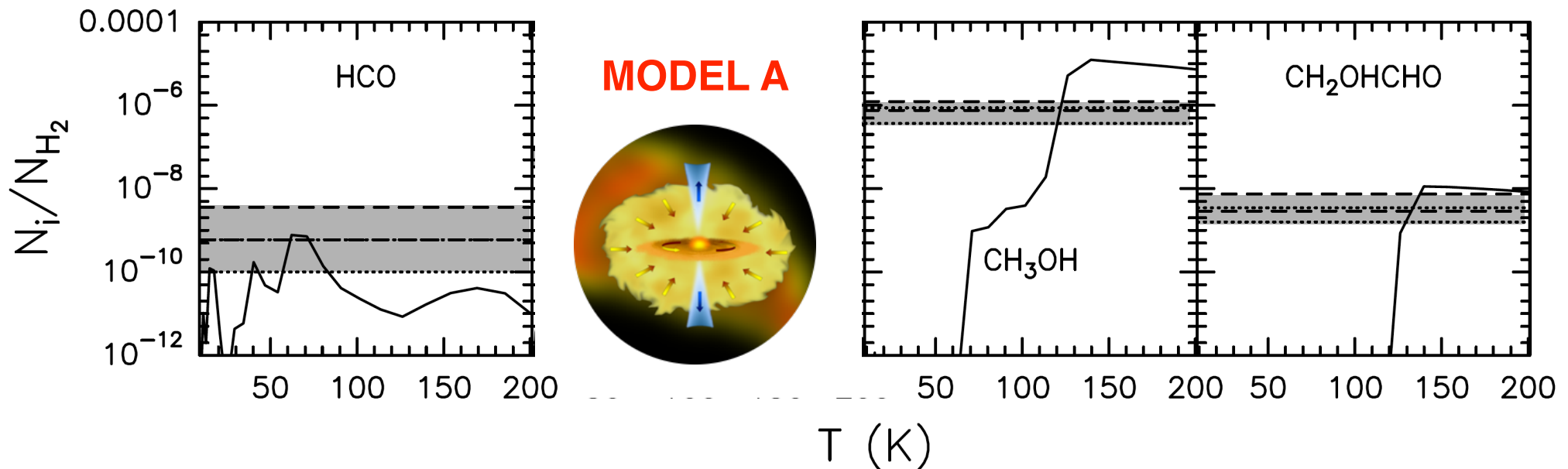
but **MODEL C** (recombination of two HCO radicals) cannot be ruled out





## The formation of glycolaldehyde $\text{CH}_2\text{OHCHO}$

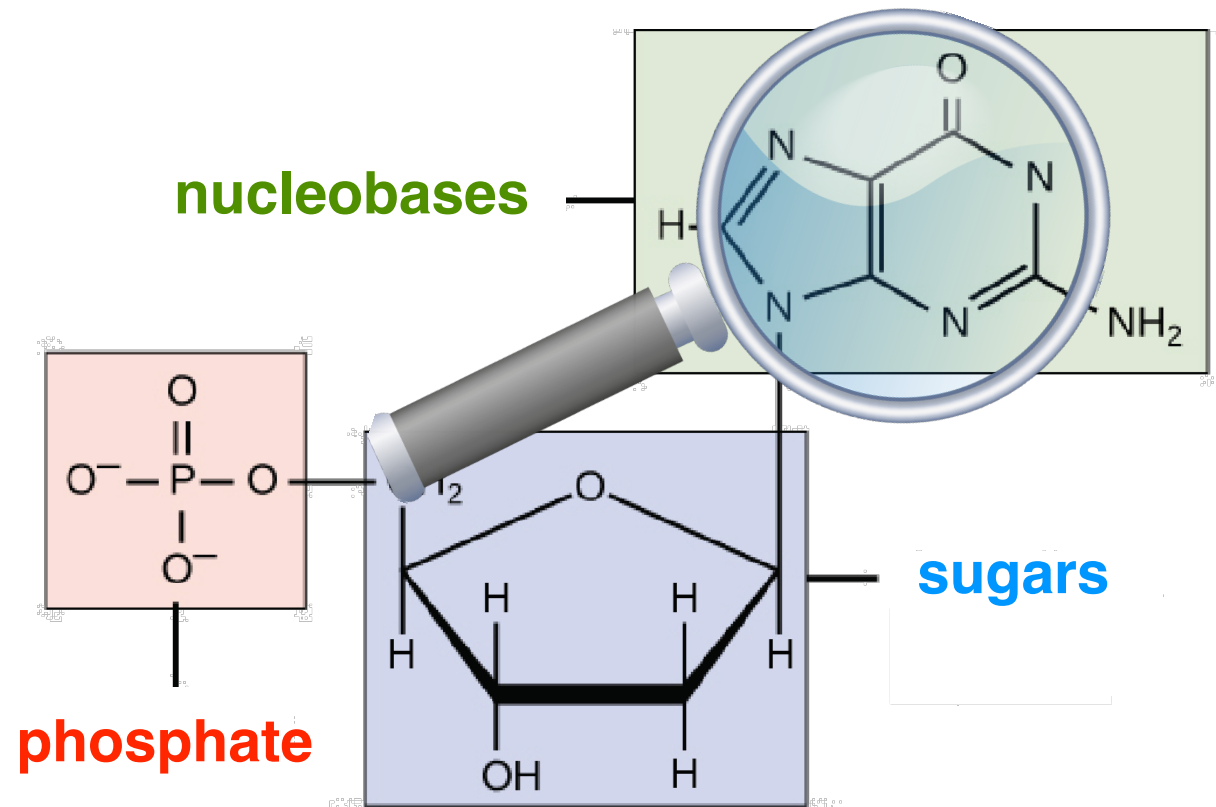
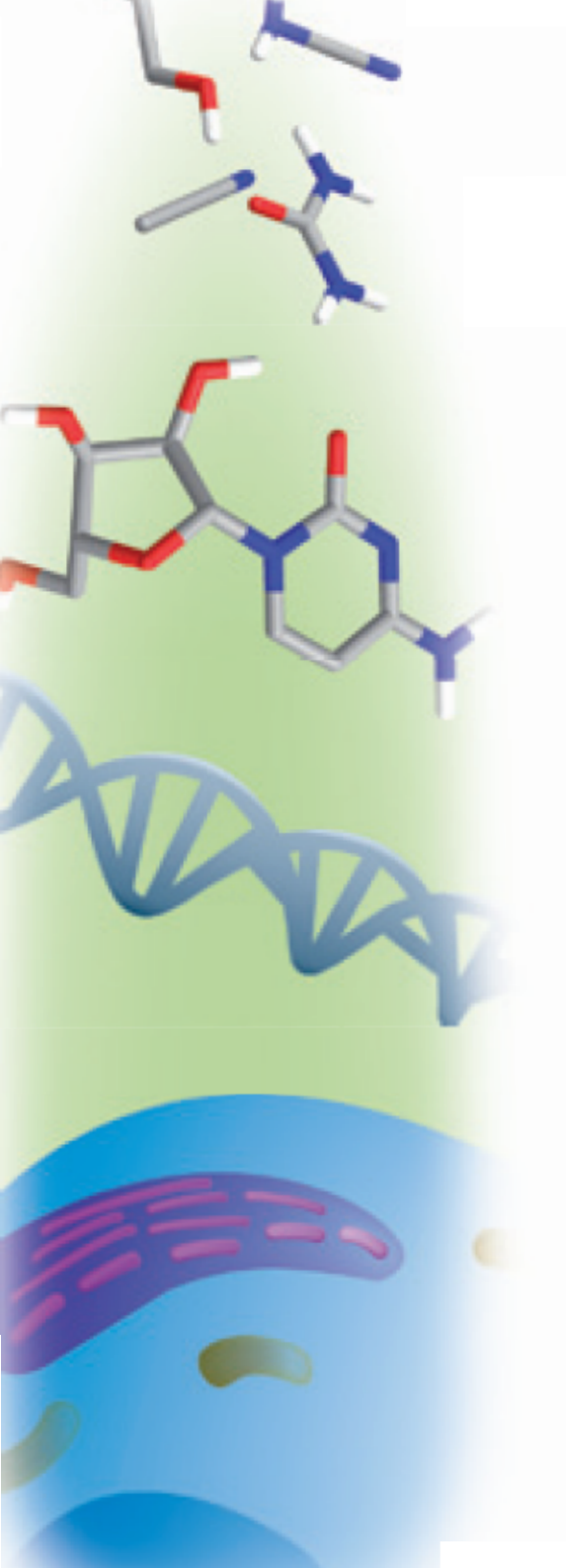
- The model that fits better the observations is **MODEL A**:



but **MODEL C** (recombination of two  $\text{HCO}$  radicals) cannot be ruled out

Good agreement with laboratory works (Fedoseev et al. 2015, Chuang et al. 2016) and chemical modeling (Coutens et al. 2018)





# BIOSFERA

Birth Of Stars and LiFE: Edge Research at INAF







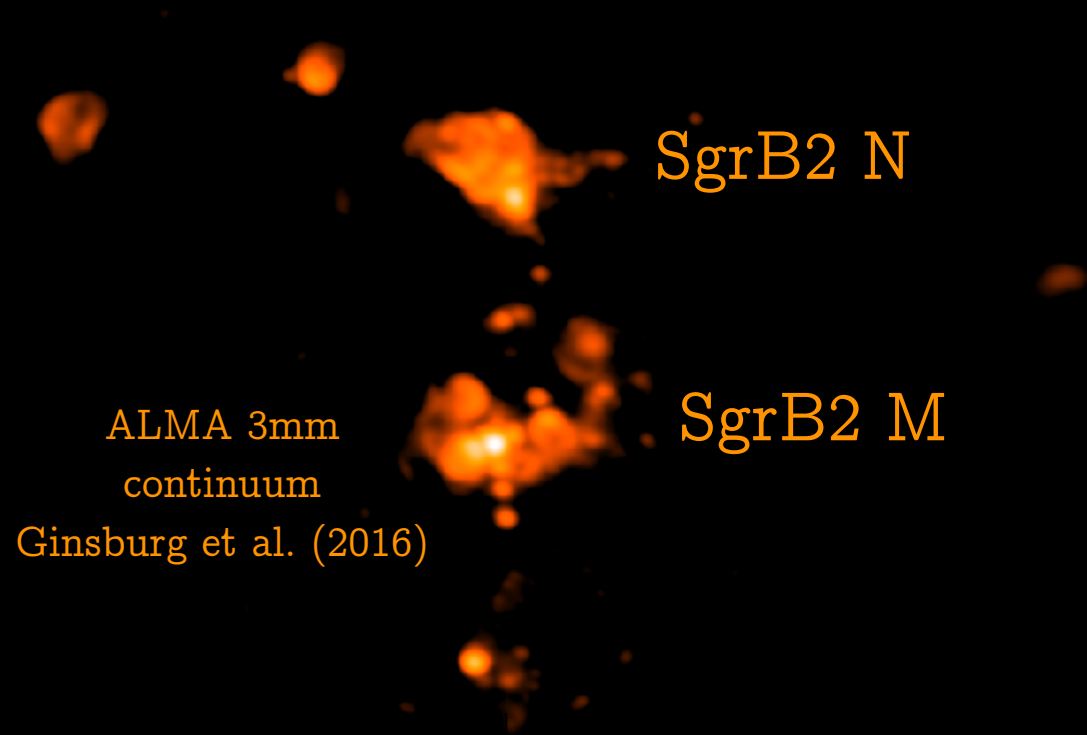
G+0.693  
molecular cloud



GALACTIC  
CENTER



# The uniqueness of G+0.693





# The uniqueness of G+0.693

G+0.693



SgrB2 N

SgrB2 M

ALMA 3mm  
continuum

Ginsburg et al. (2016)



# The uniqueness of G+0.693

C<sub>2</sub>H<sub>5</sub>OH IRAM 30m

mosaic (beam 28")

Requena-Torres, priv. comm

G+0.693

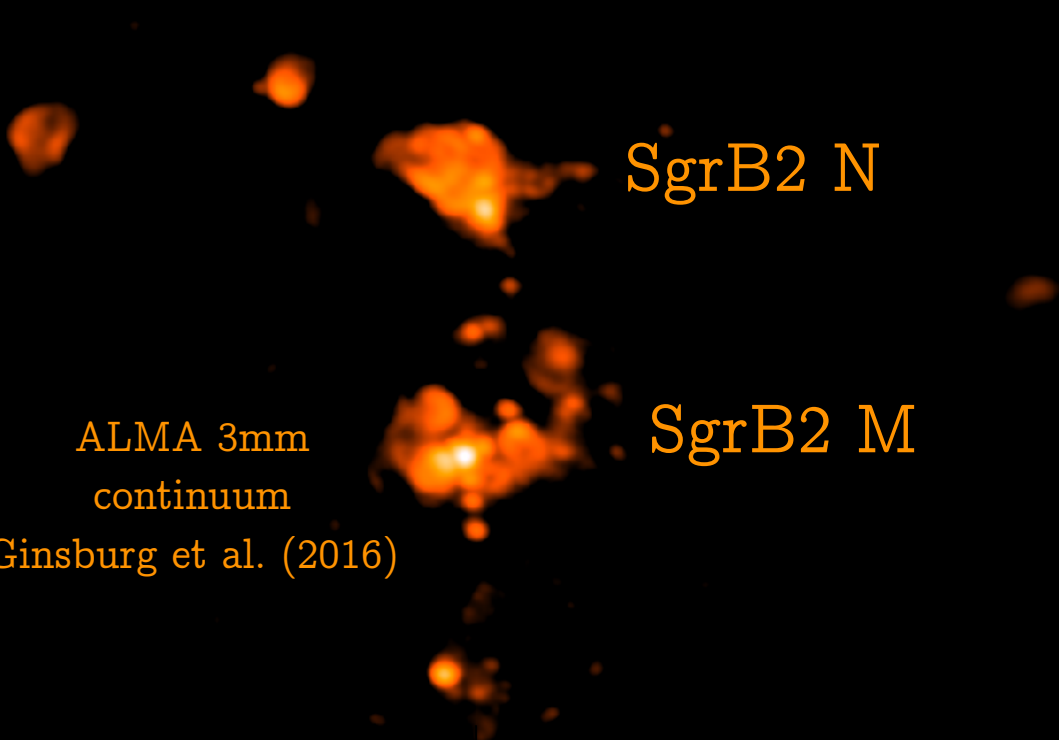


SgrB2 N

SgrB2 M

ALMA 3mm  
continuum

Ginsburg et al. (2016)



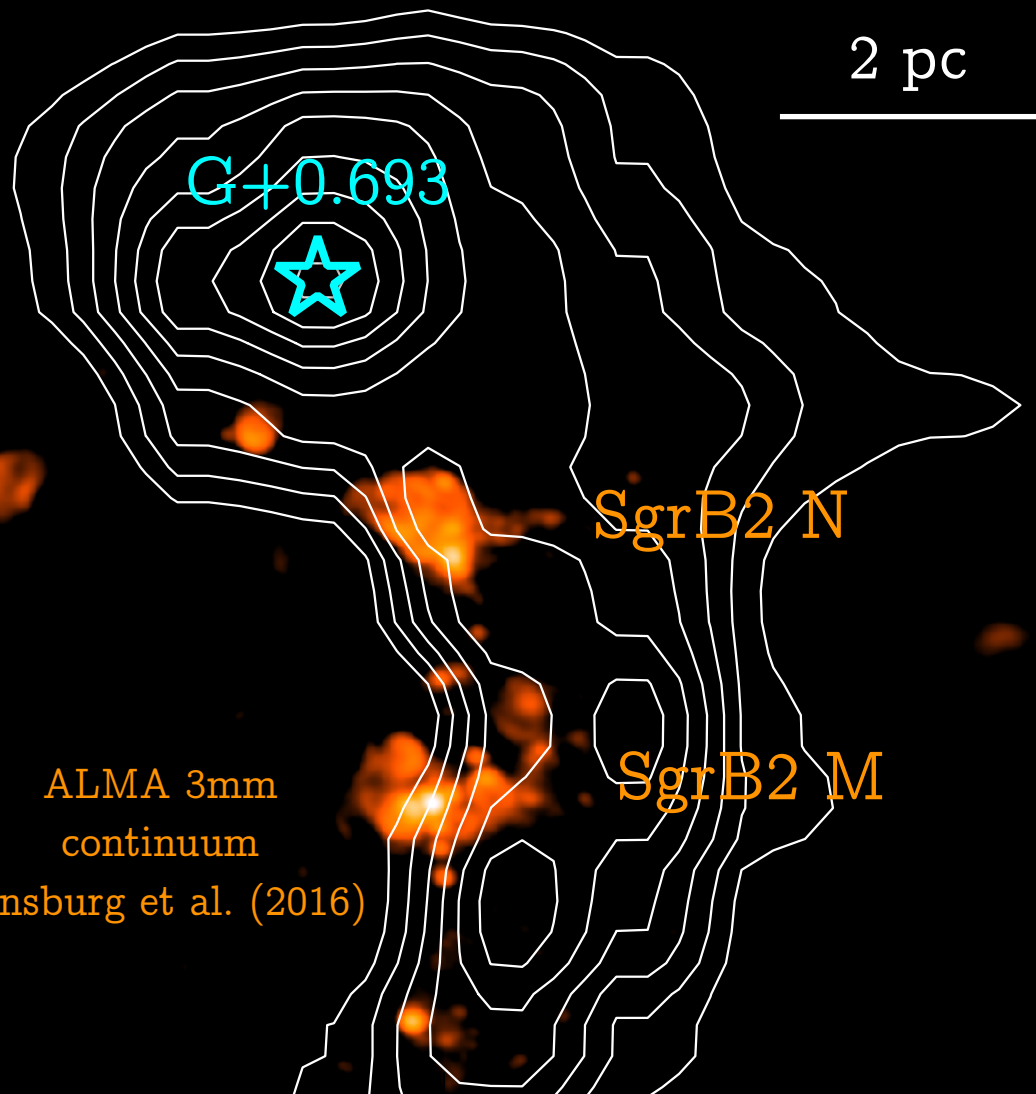


# The uniqueness of G+0.693

C<sub>2</sub>H<sub>5</sub>OH IRAM 30m

mosaic (beam 28")

Requena-Torres, priv. comm



2 pc

G+0.693

SgrB2 N

SgrB2 M

ALMA 3mm  
continuum

Ginsburg et al. (2016)



# The uniqueness of G+0.693

$\text{C}_2\text{H}_5\text{OH}$  IRAM 30m

mosaic (beam  $28''$ )

Requena-Torres, priv. comm

G+0.693



SgrB2 N

SgrB2 M

ALMA 3mm  
continuum

Ginsburg et al. (2016)

SiO(3-2)  
SMA+APEX

Zeng et al., in prep.

G+0.693



SgrB2 N





# The uniqueness of G+0.693

## SHOCKS!

C<sub>2</sub>H<sub>5</sub>OH IRAM 30m

mosaic (beam 28")

Requena-Torres, priv. comm

G+0.693



SgrB2 N

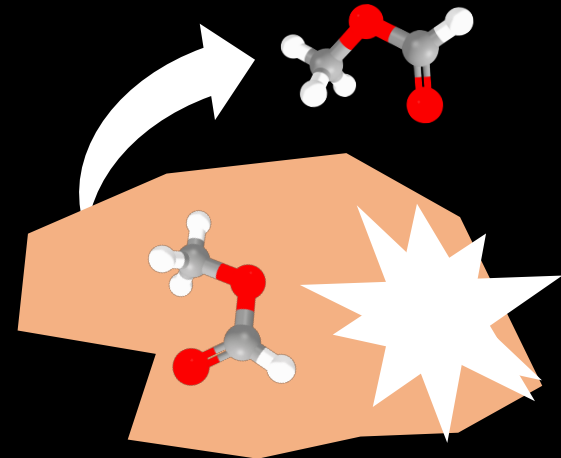
SgrB2 M

ALMA 3mm  
continuum

Ginsburg et al. (2016)

SiO(3-2)

G+0.693





# Search for cynomethanimine HNCHCN

G+0.693

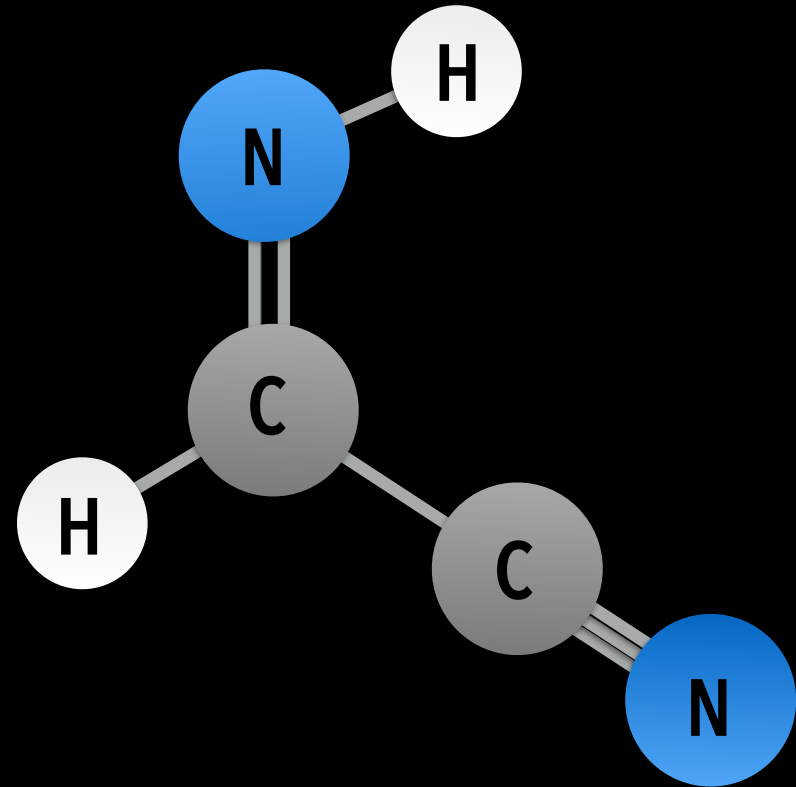


SgrB2 N

SgrB2 M

ALMA 3mm  
continuum

Ginsburg et al. (2016)





# Search for cynomethanimine HNCHCN

G+0.693

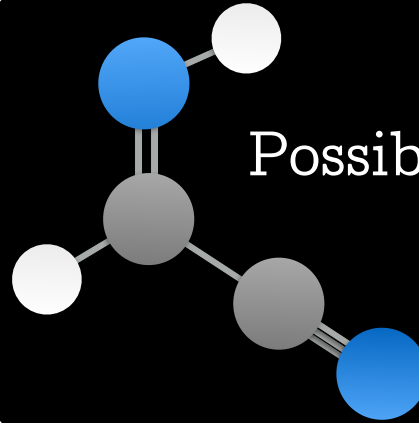


SgrB2 N

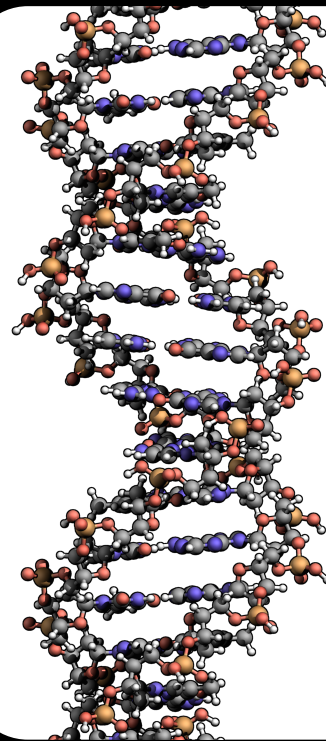
SgrB2 M

ALMA 3mm  
continuum

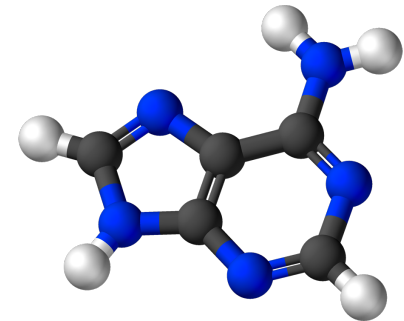
Ginsburg et al. (2016)



Possible precursor of  
adenine



Adenine



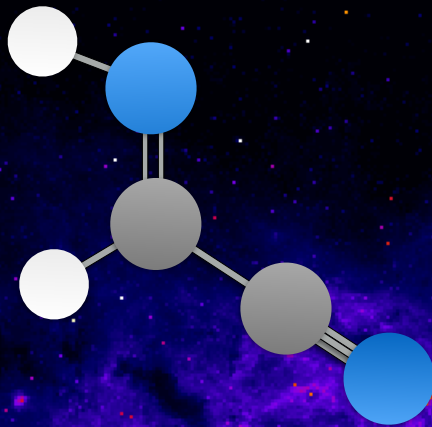
DNA and RNA  
nucleobase



# Cynomethanimine HNCHCN

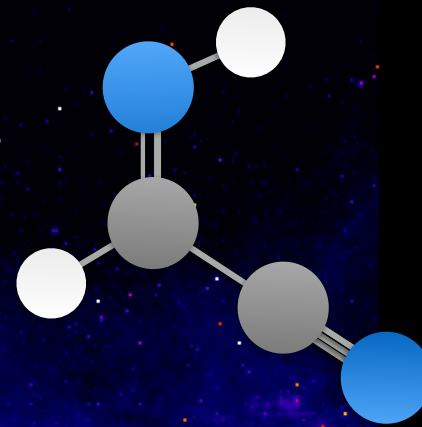
The most stable dimer of HCN is C-cyanomethanimine, which presents two isomers:

E-isomer



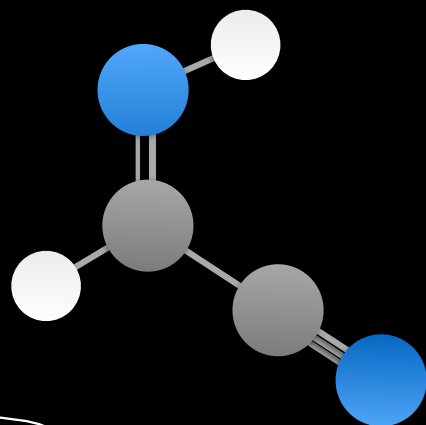
- Higher energy isomer
- Previously detected only in SgrB2 N (Zaleski et al. 2013)

Z-isomer



- Lower energy isomer
- Never detected in the ISM, despite observational efforts (Melli et al. 2018)





G+0.693





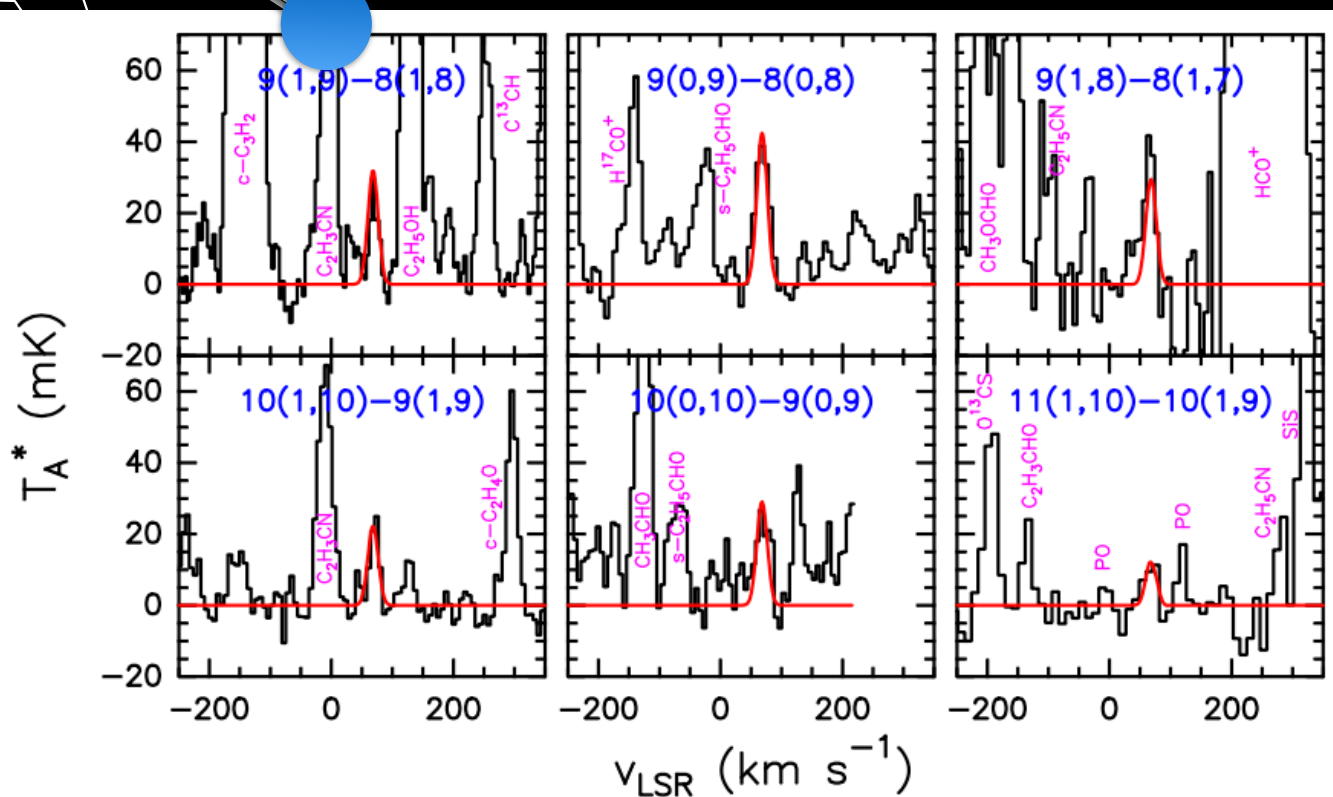
Abundant Z-cyanomethanimine in the interstellar medium:  
 paving the way to the synthesis of adenine

V. M. Rivilla<sup>1\*</sup>, J. Martín-Pintado<sup>2</sup>, I. Jiménez-Serra<sup>2</sup>, S. Zeng<sup>3</sup>, S.  
 J. Armijos-Abendaño<sup>6</sup>, M. A. Requena-Torres<sup>7</sup>, R. Aladro<sup>8</sup>, and D.

# First detection of Z-HNCHCN in the ISM

Rivilla et al. (2019b)

G+0.693





Abundant Z-cyanomethanimine in the interstellar medium:  
paving the way to the synthesis of adenine

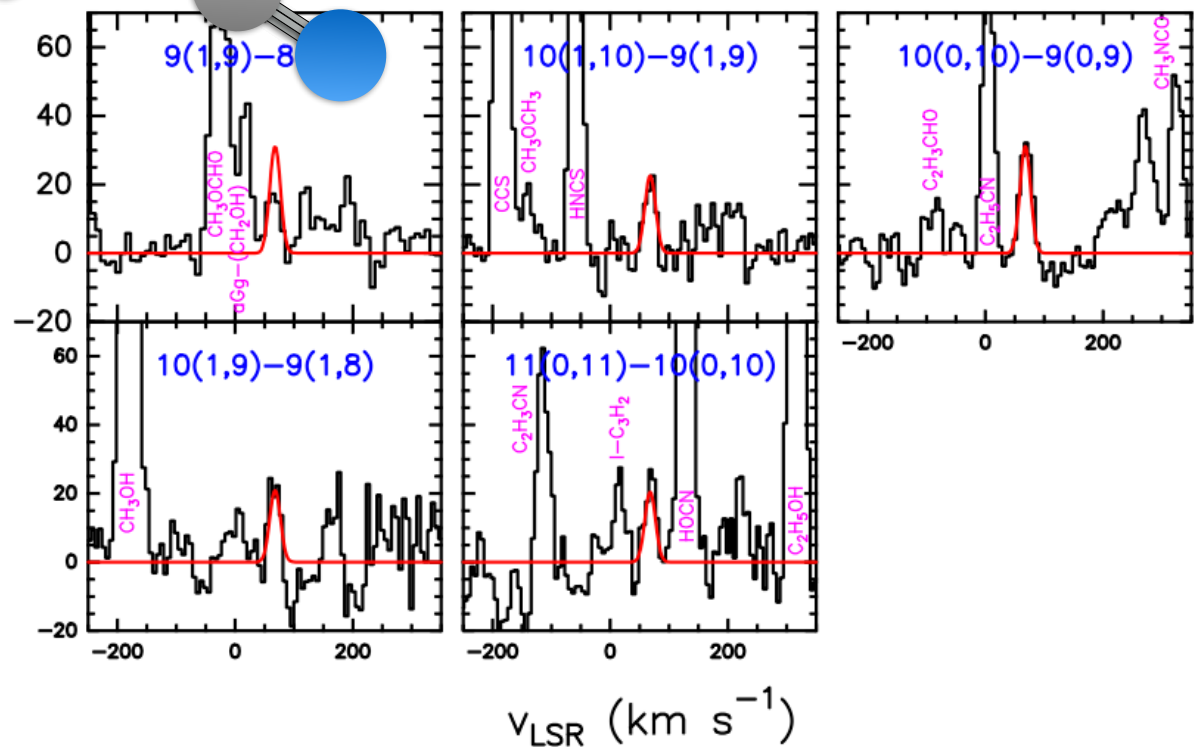
V. M. Rivilla<sup>1\*</sup>, J. Martín-Pintado<sup>2</sup>, I. Jiménez-Serra<sup>2</sup>, S. Zeng<sup>3</sup>, S. Martín<sup>4,5</sup>,  
J. Armijos-Abendaño<sup>6</sup>, M. A. Requena-Torres<sup>7</sup>, R. Aladro<sup>8</sup>, and D. Riquelme<sup>8</sup>

E-isomer also detected

Rivilla et al. (2019b)

G+0.693

$T_A^*$  (mK)

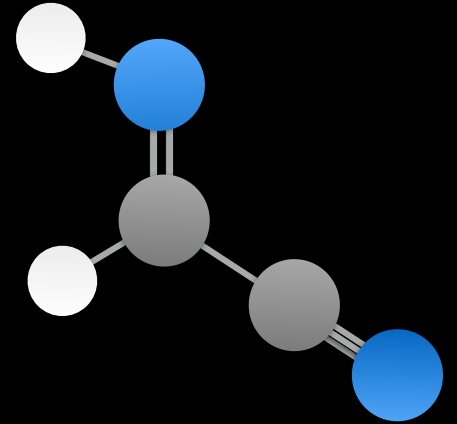
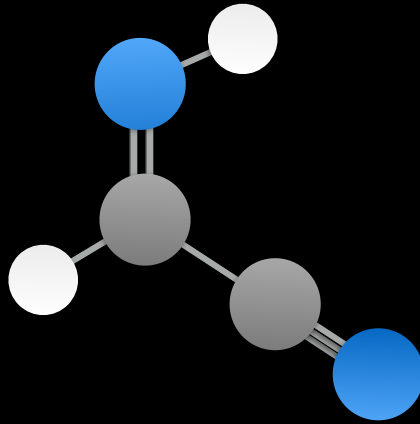




$Z/E=6$

$X(Z)=1.5 \times 10^{-9}$

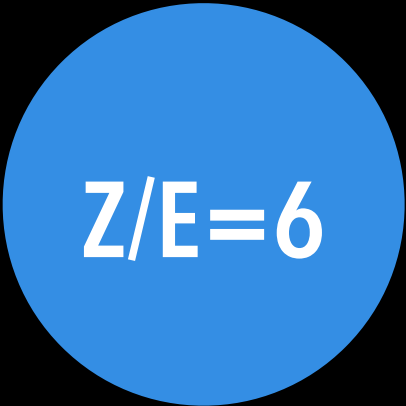
$X(E)=2.4 \times 10^{-10}$



G+0.693

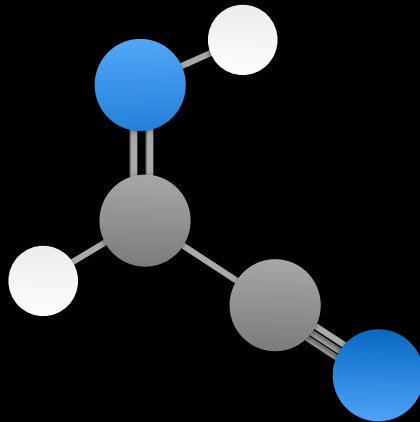




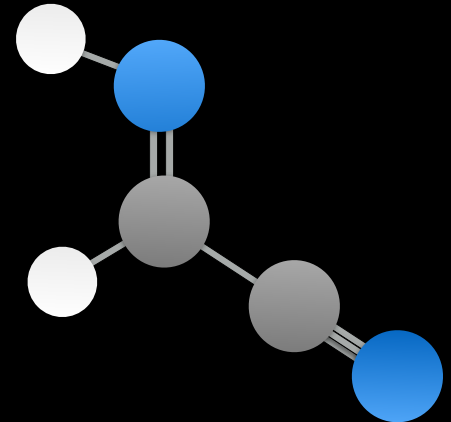


$Z/E=6$

$X(Z)=1.5 \times 10^{-9}$



$X(E)=2.4 \times 10^{-10}$



G+0.693

$X(\text{HNCHCN})=1.7 \times 10^{-9}$

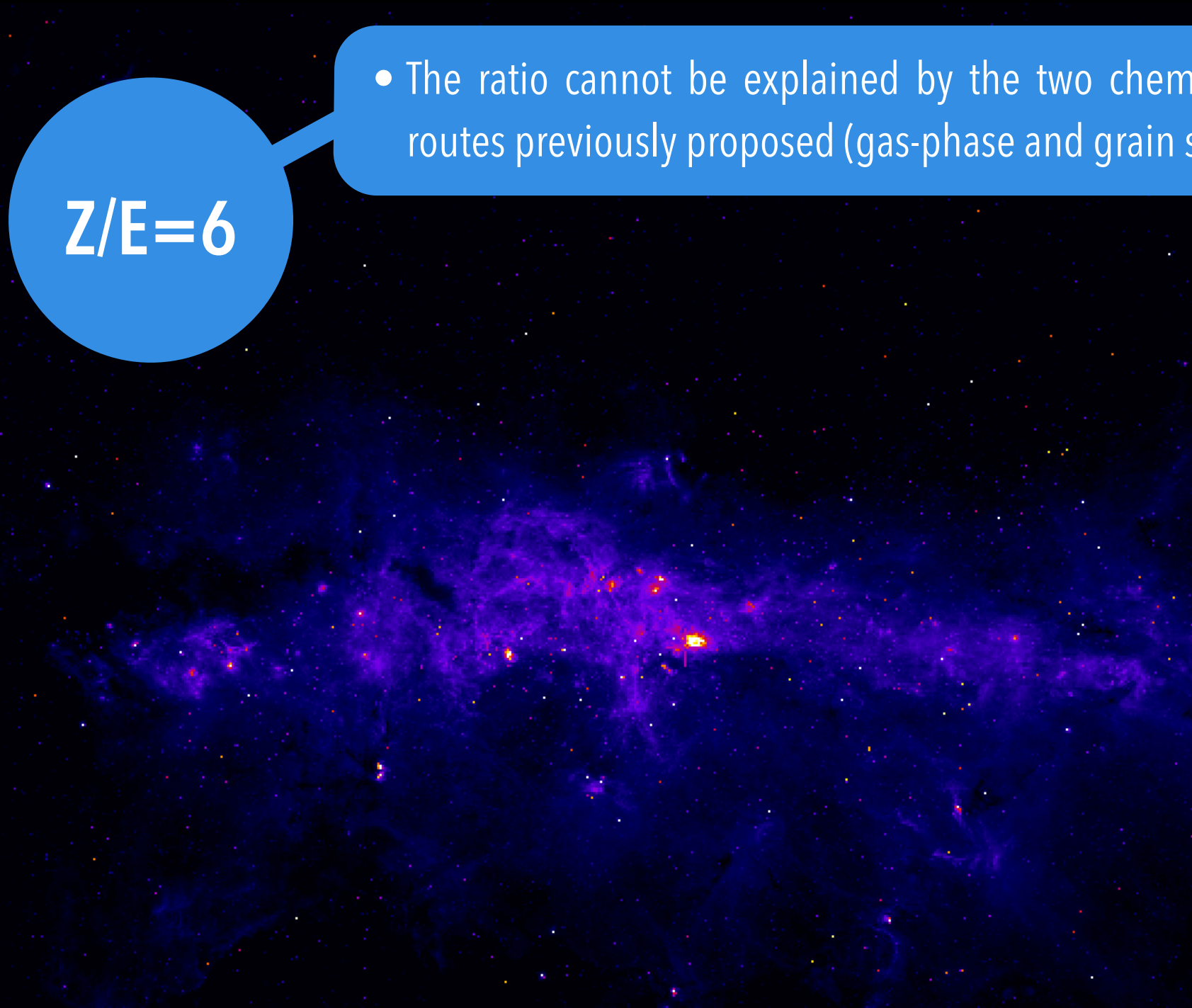
Much more abundant than  
previously thought



# Formation of HNCHCN

$Z/E=6$

- The ratio cannot be explained by the two chemical formation routes previously proposed (gas-phase and grain surface)





# Formation of HNCHCN

$Z/E=6$

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## Destruction routes?

**New ab initio calculations involving destruction reactions  
between E/Z - HNCHCN with H**

(Christopher N. Shingledecker, Germán Molpeceres and Johannes Kästner)



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- Chemical models need to be run to see if this explains  $Z/E \sim 6$



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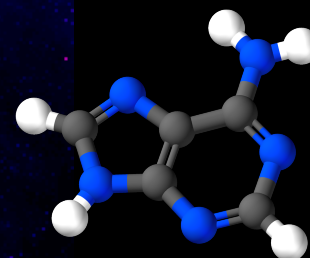
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Whatever the formation/destruction mechanism(s), the high abundance of Z-HNCHCN shows that possible precursors of adenine are efficiently formed in the ISM.

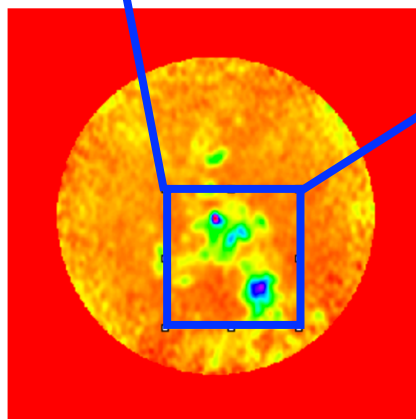
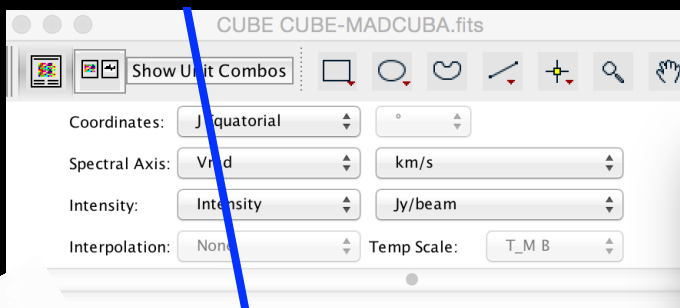
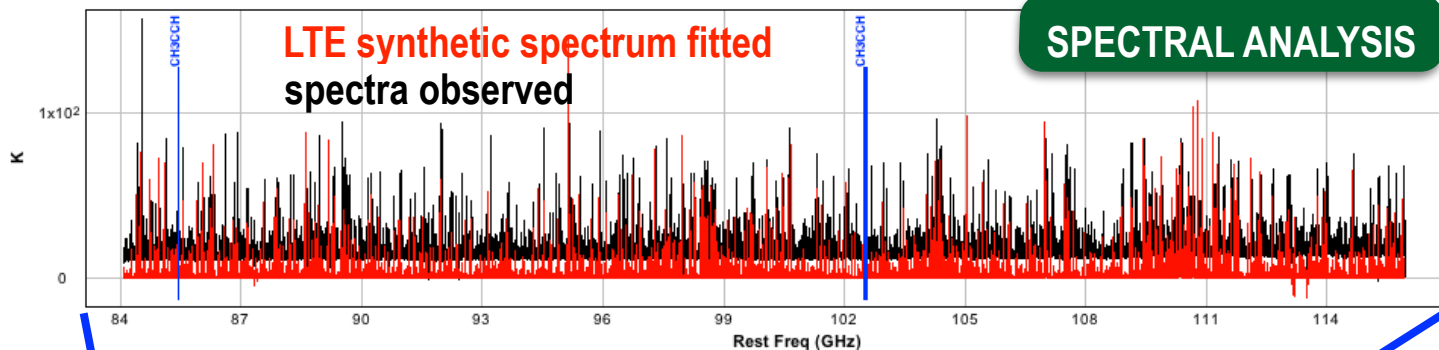




# Software development



**Spectral Line Identification and Modelling (SLIM) in the Madrid Data CUBE Analysis (MADCUBA) package**  
 An interactive software for data cube analysis  
 S. Martín<sup>1,2</sup>, J. Martín-Pintado<sup>3</sup>, C. Blanco-Sánchez<sup>3</sup>, V. M. Rivilla<sup>4</sup>, A. Rodríguez-Franco<sup>5</sup>, and F. Rico-Villas<sup>3</sup>



Pixels: X= 355 Y= 250 Z= 551.0

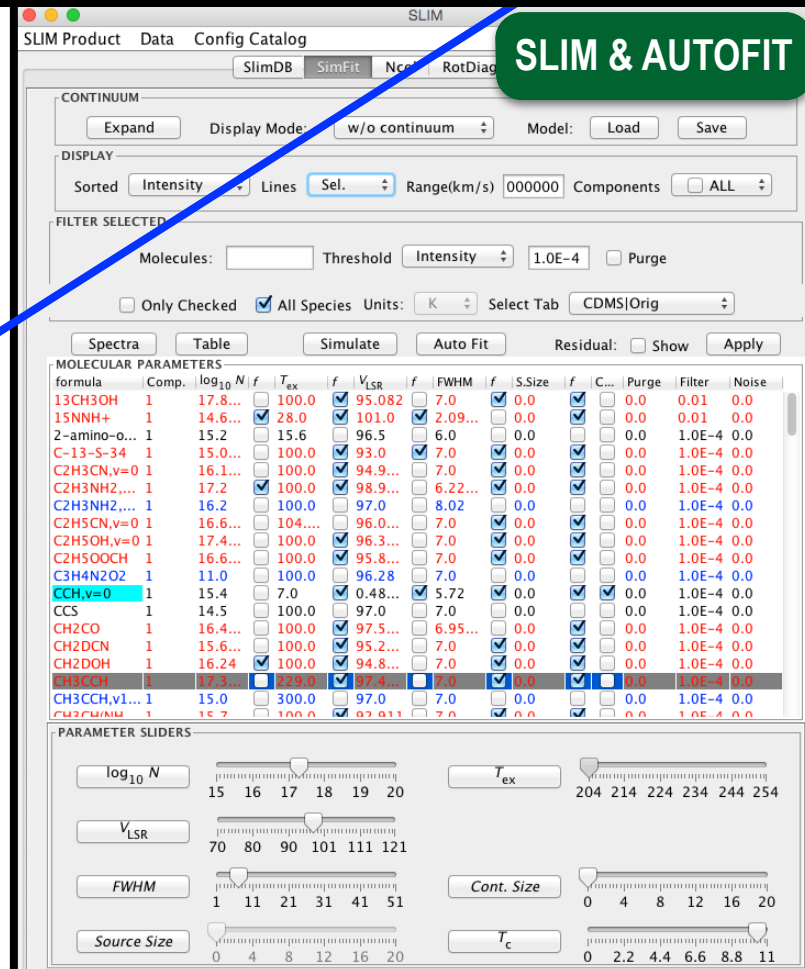
Coordinates: RA: 05h 30m 43.95161s DEC: 33° 48m 14.00587s

Spectral Axis: -4.45

Intensity:

Polarization: I

**DATAcube VISUALISATION**





- 21 refereed papers accepted (5 as first author, 5 as 2<sup>nd</sup> author)
- 2 invited talks and 8 contributed talks at conferences
- 10 observational proposals accepted as PI: 3 ALMA, 2 VLA, 4 IRAM 30m, 1 GBT



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## RESEARCH STAYS

- Max Planck Institute for Extraterrestrial Physics (MPE), Germany
- JAO ALMA Visitor Program (Chile)



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## RESEARCH STAYS

- Max Planck Institute for Extraterrestrial Physics (MPE), Germany
- JAO ALMA Visitor Program (Chile)

Co-supervisor of 2 Master's degree students and 2  
Physics degree students in the University of Florence.

## MENTORING



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## RESEARCH STAYS

- Max Planck Institute for Extraterrestrial Physics (MPE), Germany
- JAO ALMA Visitor Program (Chile)

Co-supervisor of 2 Master's degree students and 2 Physics degree students in the University of Florence.

## MENTORING

Astronomy and astrochemistry lessons to high-school students at the Osservatorio Astrofisico di Arcetri

## TEACHING



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## RESEARCH STAYS

- Max Planck Institute for Extraterrestrial Physics (MPE), Germany
- JAO ALMA Visitor Program (Chile)

Co-supervisor of 2 Master's degree students and 2 Physics degree students in the University of Florence.

## MENTORING

Astronomy and astrochemistry lessons to high-school students at the Osservatorio Astrofisico di Arcetri

## TEACHING

## PRESS RELEASES

8 Press releases in different institutions (Media INAF, ESO, ALMA, NRAO, Leiden University, Queen Mary University)



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STAYS**

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**MENTORING**

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**TEACHING****PRESS  
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**TRANSDISCIPLINARY  
IMPACT**

My research has been cited by publications in other fields (astrobiology, chemistry, spectroscopy), such as *Life*, *Nature Chemistry*, *Physical Chemistry Chemical Physics*, *Journal of Physical Chemistry*, *Chemical Communications* and *Journal of Quantitative Spectroscopy and Radiative Transfer*.

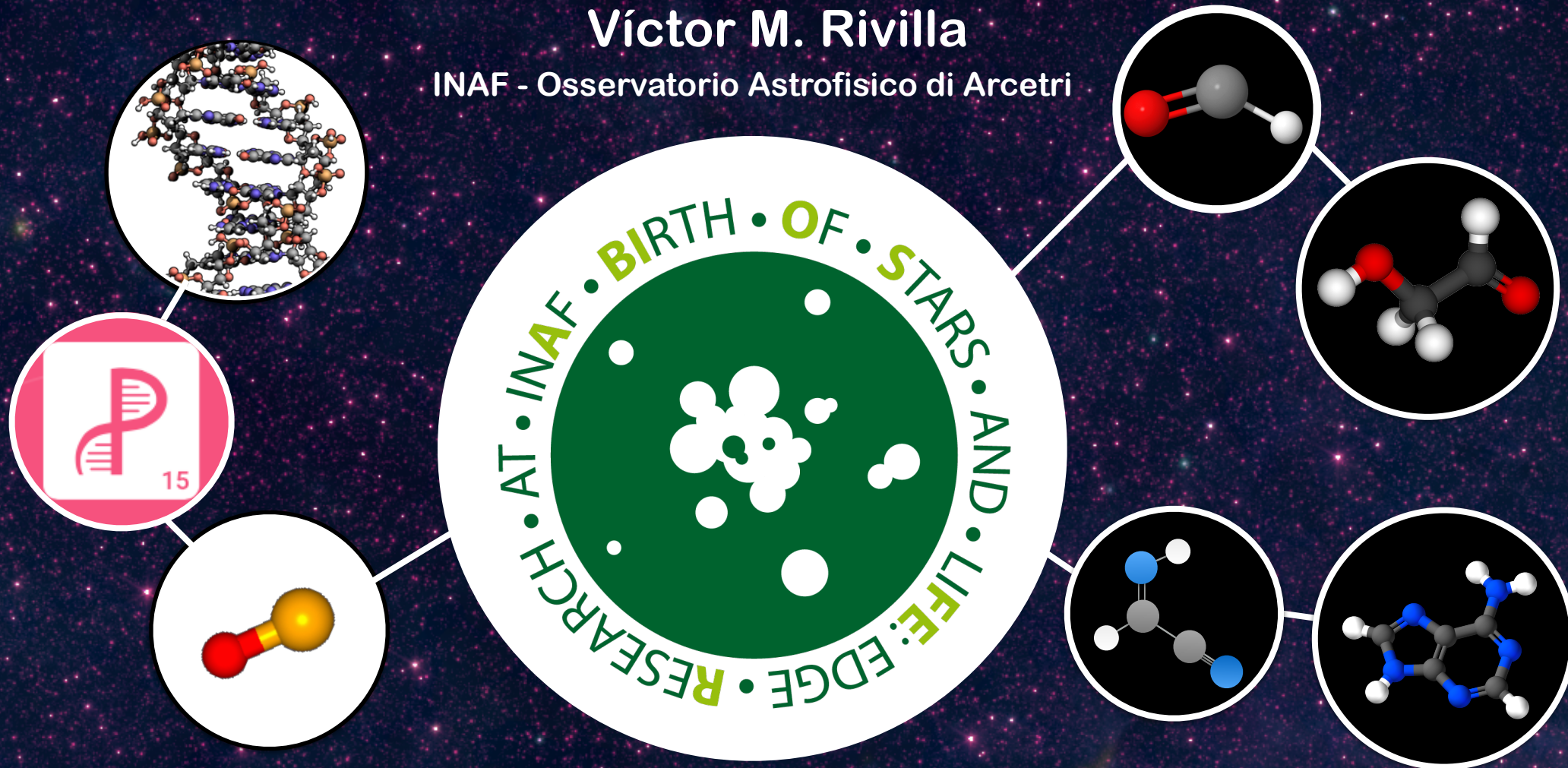


# BIOSFERA

Birth Of Stars and LiFE: Edge Research at INAF

Víctor M. Rivilla

INAF - Osservatorio Astrofisico di Arcetri



3rd ASTROFIT2 annual meeting  
Rome, Italy, October 16 2019



This project has received funding from the European Union's Horizon 2020 research and innovation programme under the Marie Skłodowska-Curie grant agreement No 664931.